

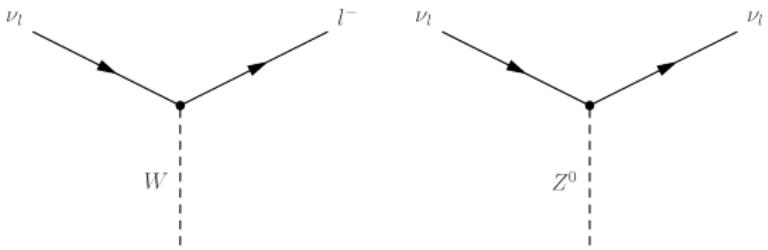
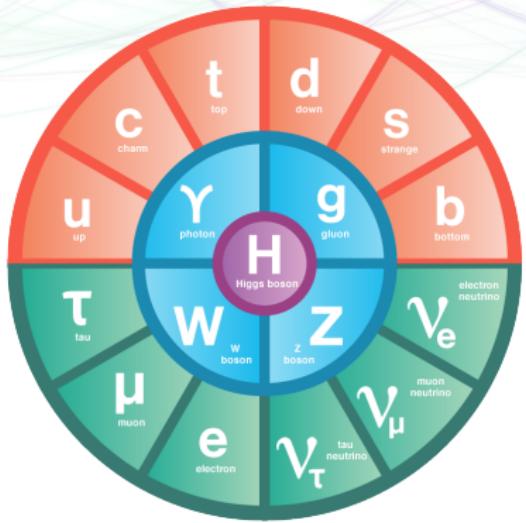
New oscillation results from NO_vA

Fermilab JETP seminar
July 20, 2016

Christopher Backhouse – Caltech
for the NO_vA collaboration



The neutrino



The neutrino

- Neutrinos mix, just like the quarks

$$|\nu_\alpha\rangle = \sum_i U_{\alpha i}^* |\nu_i\rangle$$

$$i = 1, 2, 3 \quad \alpha = e, \mu, \tau$$



- PMNS matrix. \sim CKM matrix for leptons
- Unlike the quarks, mixings are large

	CKM			PMNS		
	d	s	b	ν_e	ν_μ	ν_τ
u			.			
c			.			
t	.	.				

C. Backhouse (Caltech)

NOvA



The neutrino

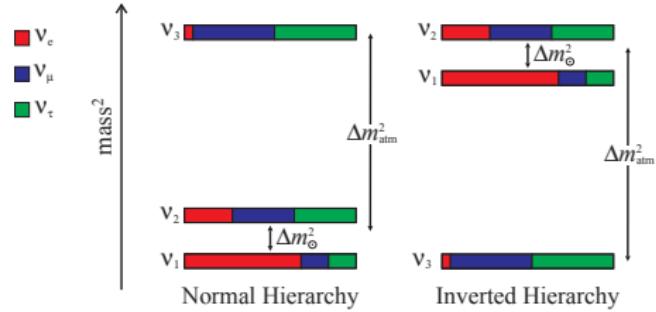
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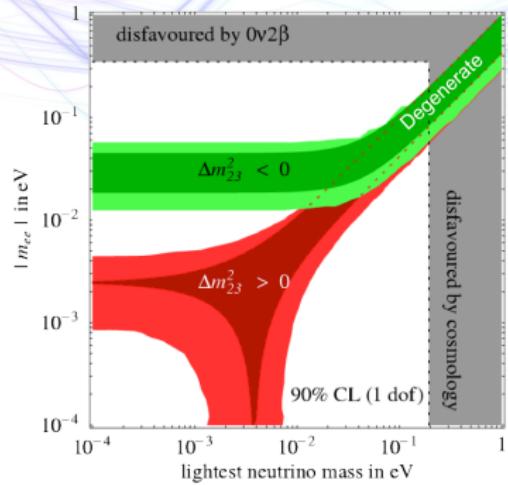
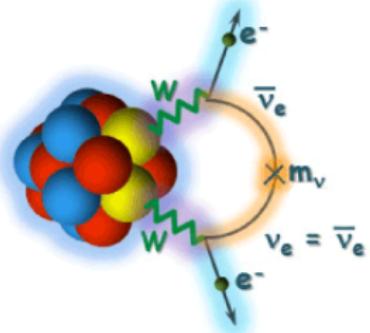


- ▶ PMNS matrix. \sim CKM matrix for leptons
- ▶ Unlike the quarks, mixings are large
- ▶ Open neutrino questions:
 - ▶ Dirac or Majorana?
 - ▶ Absolute masses
 - ▶ Ordering of the mass states
 - ▶ CP -violation?
 - ▶ Random mixing parameters, or patterns?



Why hierarchy?

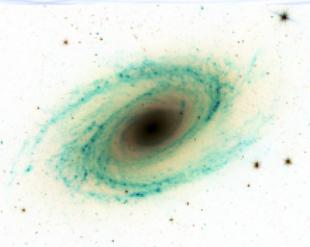
- ▶ Is the most electron-like state lightest?
- ▶ i.e. Does the pattern of the masses match the charged leptons?



- ▶ Are neutrinos Majorana particles ($\nu = \bar{\nu}$)?
- ▶ Observation of $0\nu\beta\beta$ would be proof they are
- ▶ Impact of **IH** determination: lack of $0\nu\beta\beta$ implies Dirac nature

Why CPV?

- ▶ Does e.g. $P(\nu_\mu \rightarrow \nu_e) = P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$?
- ▶ Insight into fundamental symmetries of the lepton sector
- ▶ Why is the universe not equal parts matter and antimatter?
- ▶ Sakharov conditions
 - ▶ Baryon number violation
 - ▶ Out of thermal equilibrium
 - ▶ C and CP violation
- ▶ CPV in the Standard Model, eg for K and B mesons, but too small
- ▶ “Leptogenesis”: generate asymmetry in neutrinos, transfer to baryons
- ▶ Require neutrino **appearance** experiment to discover



Mixing patterns

ν_3



- ▶ Only a small fraction of ν_e in $|\nu_3\rangle$ (the famous $\sin^2 2\theta_{13}$)
- ▶ The remainder is split $\sim 50/50 \nu_\mu/\nu_\tau$ ($\sin^2 \theta_{23}$)
- ▶ Accident? Or a sign of underlying structure?

- ▶ Is θ_{23} exactly 45° ?
- ▶ If not, is it...
 - ▶ $< 45^\circ$ ($|\nu_3\rangle$ more ν_τ , like the quarks)
 - ▶ $> 45^\circ$ ($|\nu_3\rangle$ more ν_μ , unlike quarks)

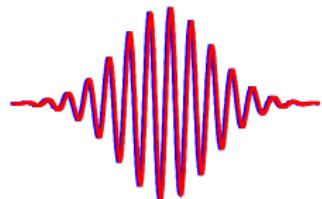
Neutrino oscillations



$$|\nu_\alpha\rangle = \sum_i U_{\alpha i}^* |\nu_i\rangle$$

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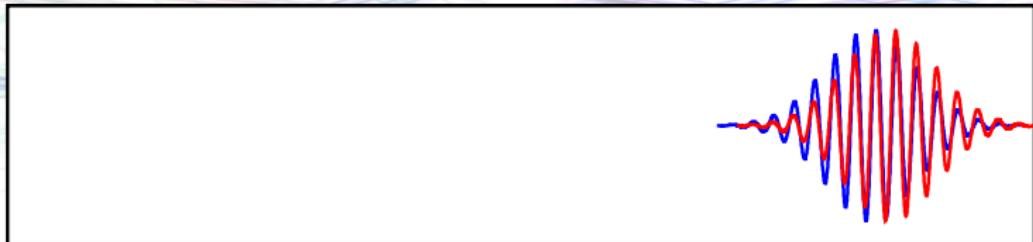
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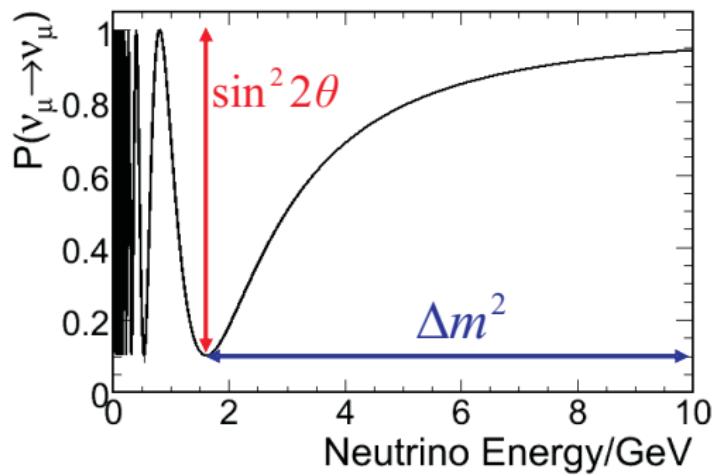
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Neutrino oscillations



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$$P_{\alpha\beta} = \left| \sum_i U_{\alpha i}^* e^{-im_i^2 L/2E} U_{\beta i} \right|^2$$



Three flavors

$$U = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$$\theta_{23} \sim 45^\circ$$

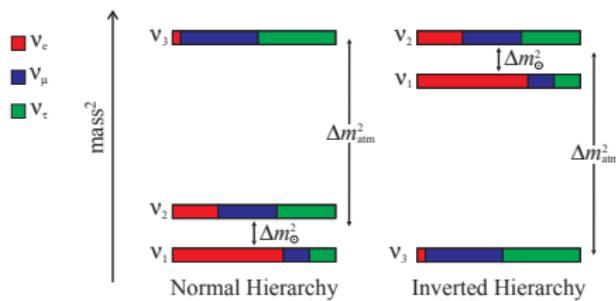
$$\Delta m_{32}^2 \sim \pm 2.5 \times 10^{-3} \text{ eV}^2$$

$$\delta_{CP} = ?$$

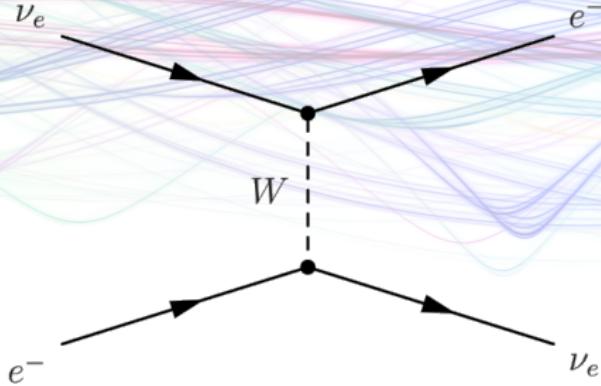
$$\theta_{13} \sim 8.5^\circ$$

$$\theta_{12} \sim 33^\circ$$

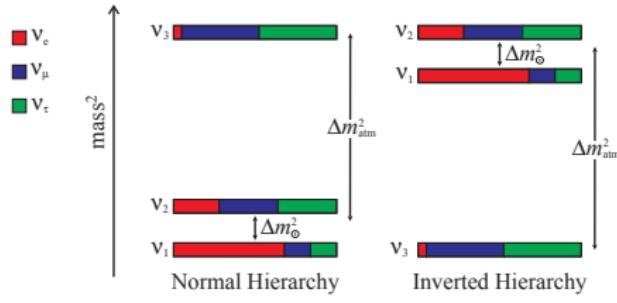
$$\Delta m_{21}^2 \sim 7.5 \times 10^{-5} \text{ eV}^2$$



Matter effects



- ▶ Electrons in the Earth drag on the “electron” neutrino states
- ▶ Sign of the effect opposite for antineutrinos and for NH/IH



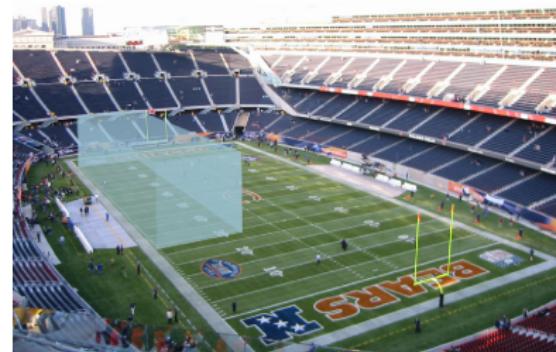
NOvA

- ▶ Requirements for neutrino oscillation experiment
 - ▶ High power neutrino source
 - ▶ Large detector
 - ▶ Good resolution of signal from background
 - ▶ Good control of systematic uncertainties



NOvA

- ▶ Requirements for neutrino oscillation experiment
 - ▶ High power neutrino source
 - ▶ Large detector
 - ▶ Good resolution of signal from background
 - ▶ Good control of systematic uncertainties
- ▶ For hierarchy and CPV
 - ▶ Appearance mode
 - ▶ Long baseline
 - ▶ Ability to study neutrinos and antineutrinos



The NOvA collaboration

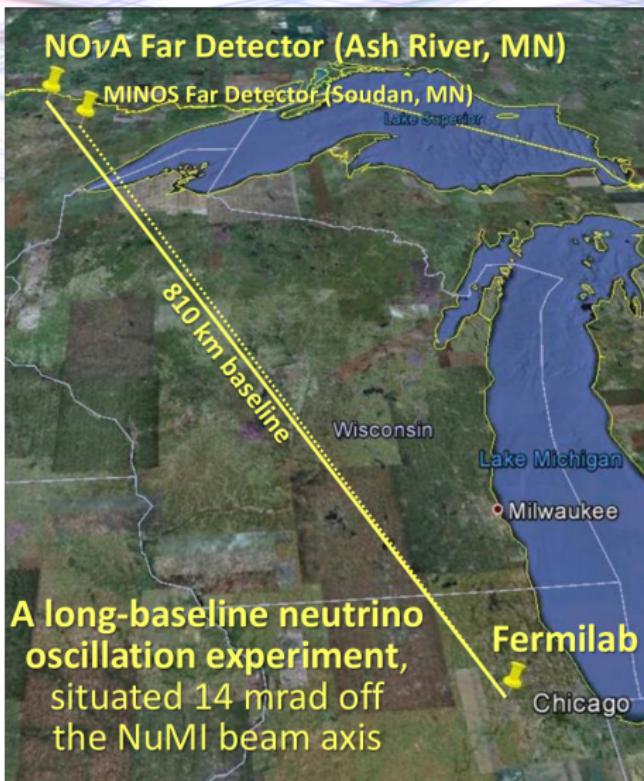


41 institutions, 7 countries, over 200 collaborators

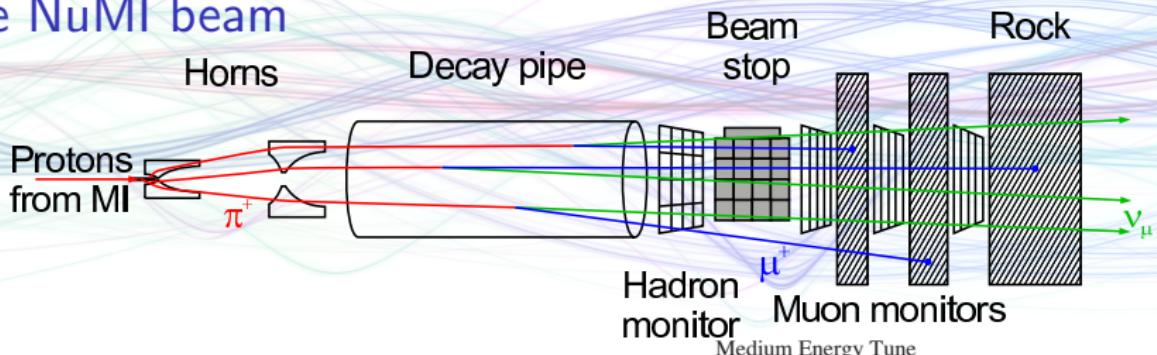
Argonne, Atlantico, Banaras Hindu, Caltech, CUSAT, Czech Academy of Sciences, Charles, Cincinnati, Colorado State, Czech Technical University, Delhi, Dubna, Fermilab, Goias, IIT-Guwahati, Harvard, IIT-Hyderabad, Hyderabad, Indiana, Iowa State, Jammu, Lebedev, Michigan State, Minnesota-Twin Cities, Minnesota-Duluth, INR Moscow, Panjab, SDMT, South Carolina, SMU, Stanford, Sussex, Tennessee, Texas-Austin, Tufts, UCL, Virginia, Wichita State, William and Mary, Winona State.

NOvA 10,000ft view

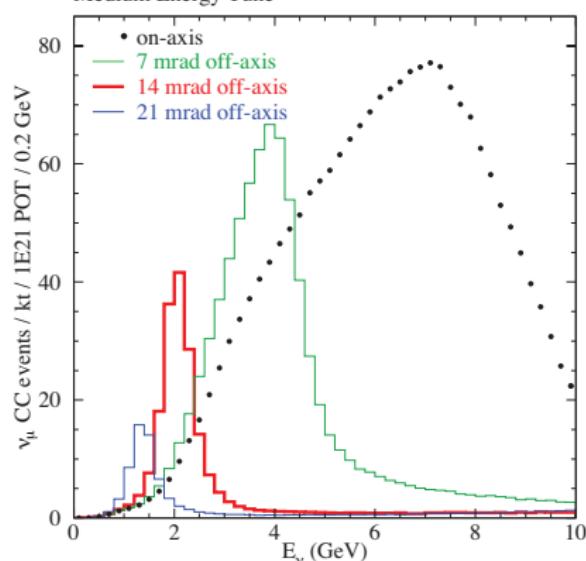
- ▶ ν_μ beam from Fermilab
- ▶ Detector 810km away in MN
- ▶ Smaller detector onsite to measure flux before oscillations
 - ▶ $\nu_\mu \rightarrow \nu_\mu$
 - ▶ $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$
 - ▶ $\nu_\mu \rightarrow \nu_e$
 - ▶ $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$
- ▶ Precision measurements of $|\Delta m_{32}^2|$ and θ_{23}
- ▶ Determine the mass hierarchy
- ▶ Search for $\delta_{CP} \neq 0$
- ▶ Plus: xsec in the ND, exotics
- ▶ Sterile neutrinos, next week!



The NuMI beam

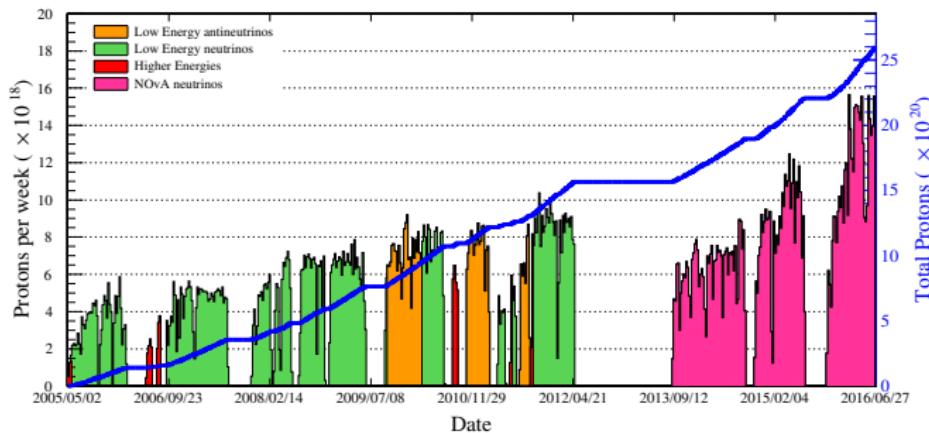


- ▶ 120 GeV protons from Main Injector
- ▶ Produce mainly π^\pm and K^\pm
- ▶ Focused by two magnetic horns
- ▶ Allow us to select charge sign for a neutrino or antineutrino beam
- ▶ 14.6mrad off-axis: neutrino energy \sim independent of pion energy
- ▶ Beam peak from 1-3GeV

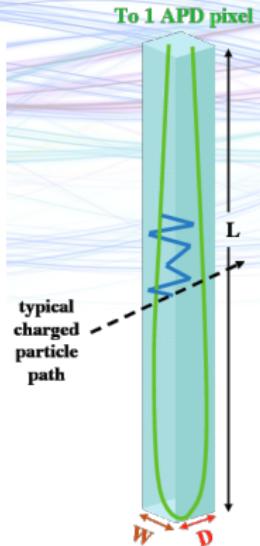
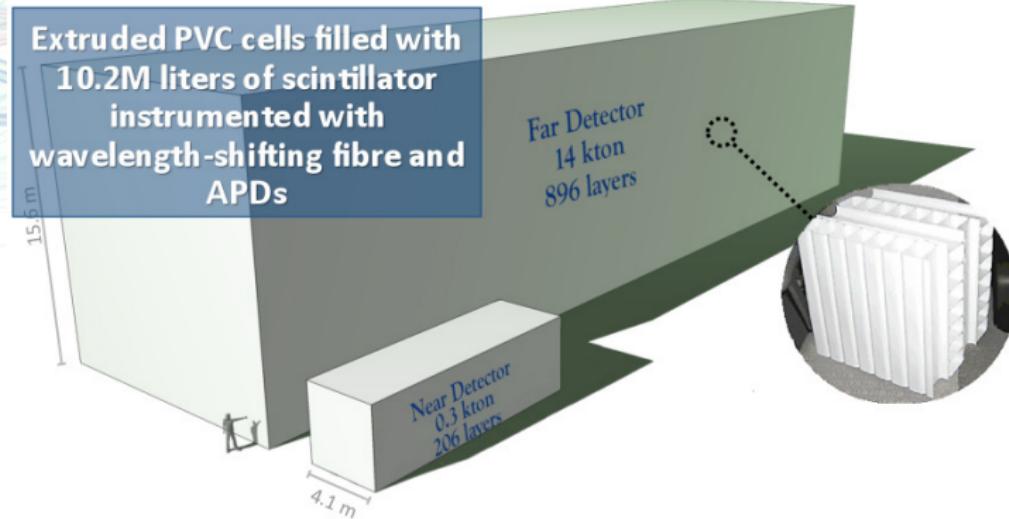


NuMI performance

- ▶ Fantastic beam performance
- ▶ Currently running at 560kW
- ▶ Achieved 700kW design intensity (briefly) in tests on June 13
- ▶ Analysis uses data from Feb 6 2014 to May 2 2016
- ▶ Equivalent to 6.05×10^{20} POT in full 14 kton detector (1 TDR year)



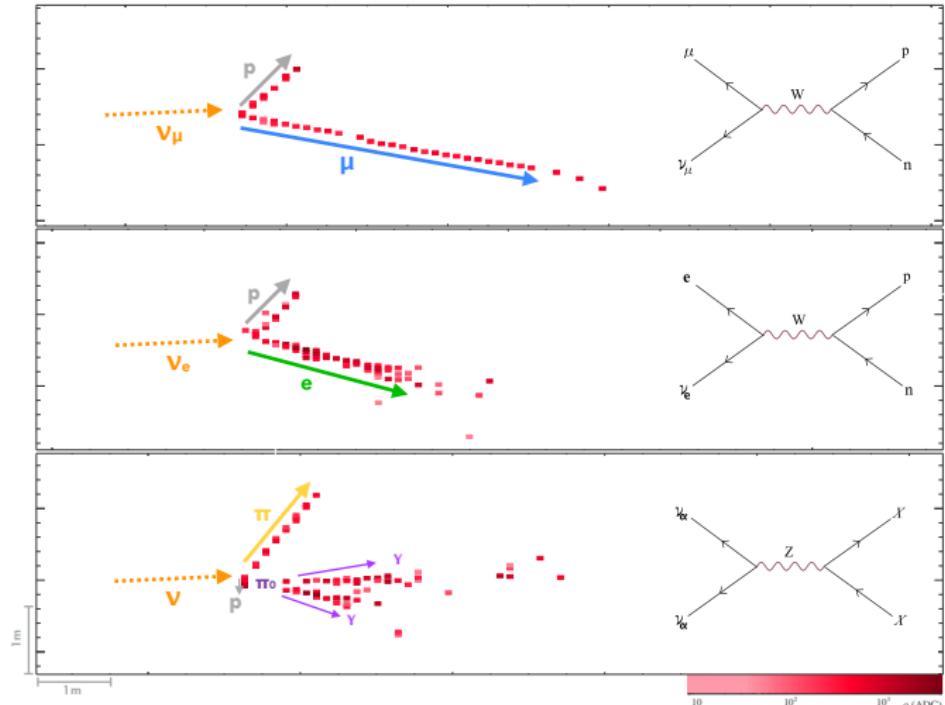
Detector technology



- ▶ Fine-grained low-Z, highly active, tracking calorimeter
- ▶ 64% liquid scintillator by mass
- ▶ 344,000 channels in 14 kton FD, on surface
- ▶ 300 ton ND, underground at FNAL



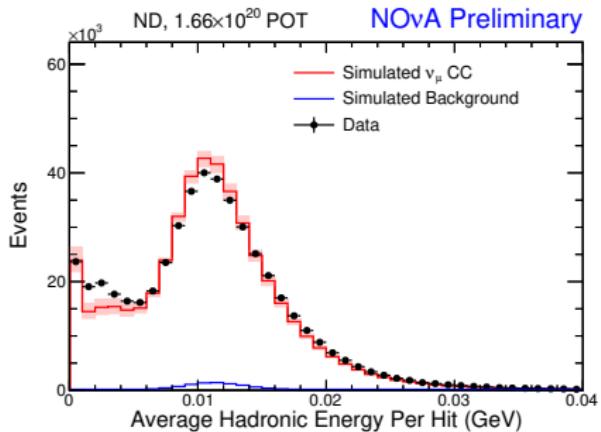
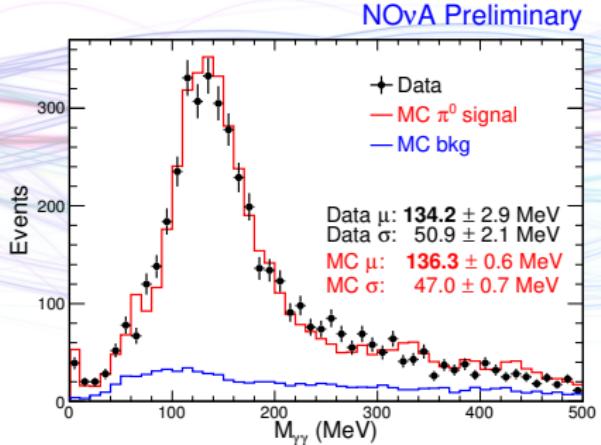
Event topologies



- ▶ Very good granularity, especially considering scale
- ▶ $X_0 = 38\text{cm}$ (6 cell depths, 10 cell widths)

Calibration and energy scale

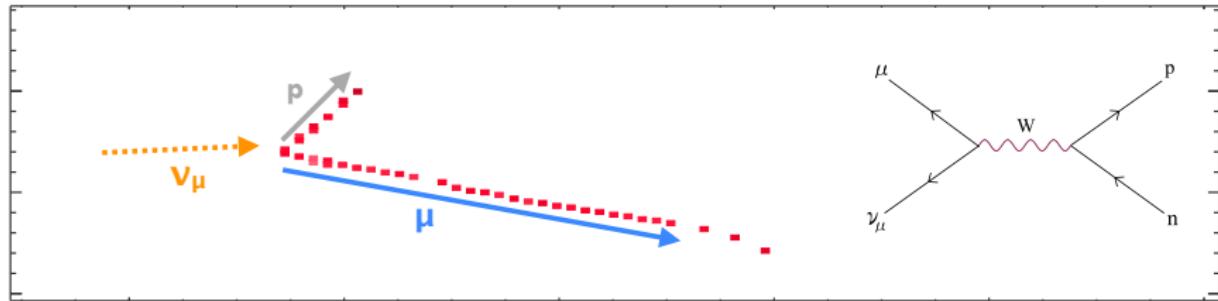
- ▶ Response varies substantially along cell due to light atten.
- ▶ Use cosmic ray muons as a standard candle to calibrate every channel individually
- ▶ Use dE/dx near the end of stopping muon to set abs. scale
- ▶ Multiple calibration x-checks
 - ▶ Beam muon dE/dx
 - ▶ Michel energy spectrum
 - ▶ π^0 mass peak
 - ▶ Hadronic energy/hit
- ▶ Take 5% abs. and rel. errors on energy scale



Principle of the ν_μ measurement



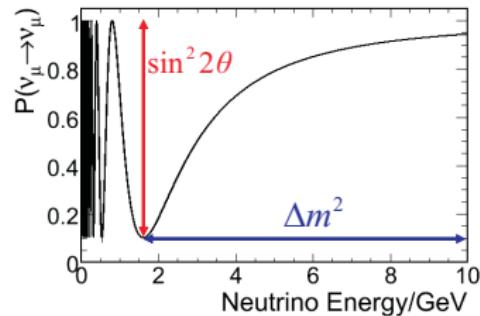
- ▶ Separate ν_μ CC interactions from backgrounds
 - ▶ Long muon track with distinctive dE/dx easy to spot
- ▶ Extrapolate observed ND spectrum to make FD unosc. prediction
- ▶ Measure shape of ν_μ deficit in the FD



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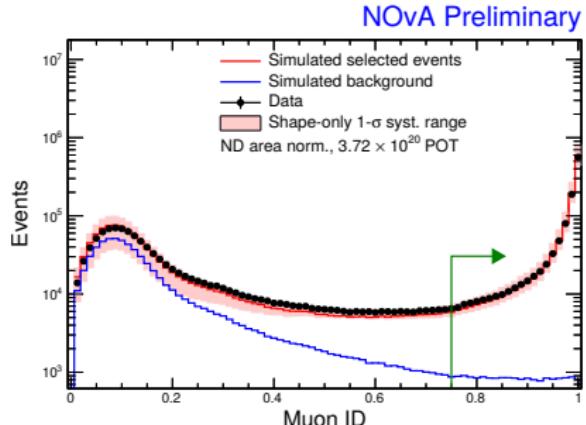
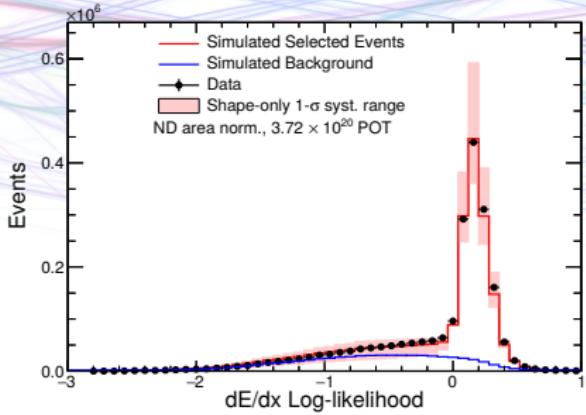


- ▶ Separate ν_μ CC interactions from backgrounds
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- ▶ Extrapolate observed ND spectrum to make FD unosc. prediction
- ▶ Measure shape of ν_μ deficit in the FD
- ▶ Two flavor approx. works well here
- ▶ $P_{\mu\mu} \approx 1 - \sin^2 2\theta_{23} \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E} \right)$
- ▶ $\theta_{23} \approx 45^\circ \rightarrow$ almost all ν_μ expected to disappear at oscillation max.

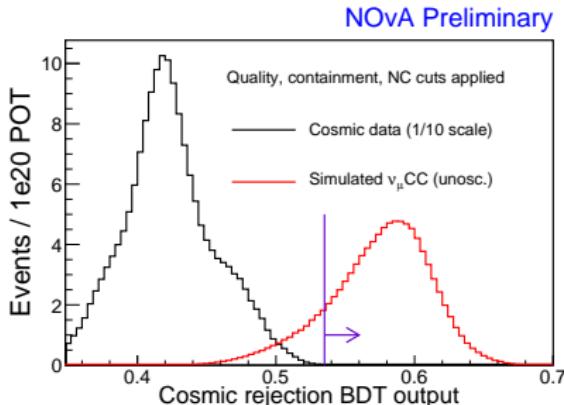
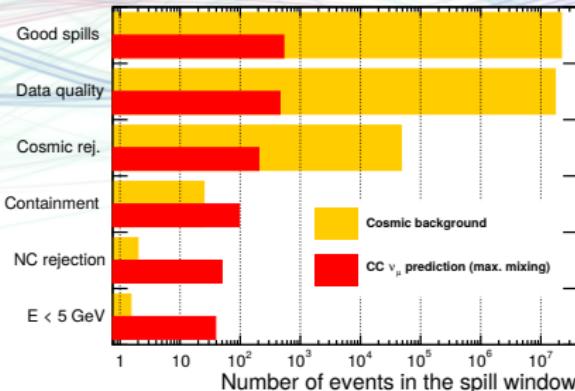


Selecting muon neutrinos

- ▶ Basic containment cuts requiring no activity close to detector walls
- ▶ kNN-based ν_μ classifier using 4 inputs
 - ▶ Track length
 - ▶ dE/dx
 - ▶ Scattering
 - ▶ Fraction of planes that have track-only
- ▶ Selection 81% efficient for ν_μ signal, 95% pure

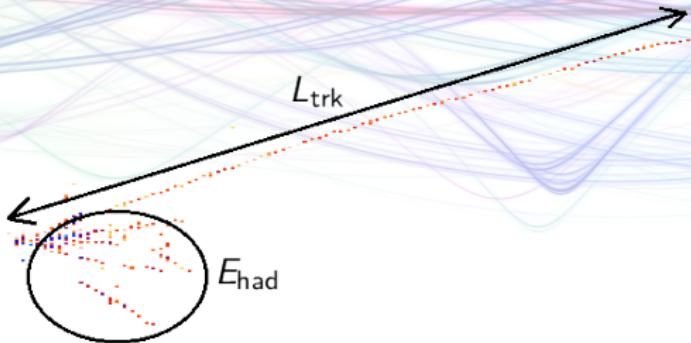


Cosmic rejection for ν_μ analysis



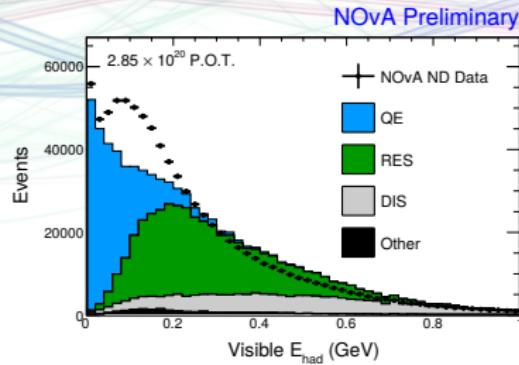
- ▶ $10\mu\text{s}$ spill window at $\sim 1\text{Hz}$ gives 10^5 rejection
- ▶ Cosmic background rate measured from data adjacent in time to the beam spill window
- ▶ Additional factor 10^7 from event topology plus boosted decision tree based on
 - ▶ Track direction
 - ▶ Track start and end points
 - ▶ Track length
 - ▶ Energy
 - ▶ Number of hits

Muon neutrino energy reconstruction



- ▶ Estimate energy of selected events to trace out oscillation structure
- ▶ Known muon $dE/dx \rightarrow E_\mu = f(L_{\text{trk}}) \sim k \times L_{\text{trk}}$
- ▶ Hadronic part of the event estimated calorimetrically
- ▶ $E_\nu = f(L_{\text{trk}}) + E_{\text{had}}$
- ▶ Achieve 7% energy resolution

Nuclear correlations



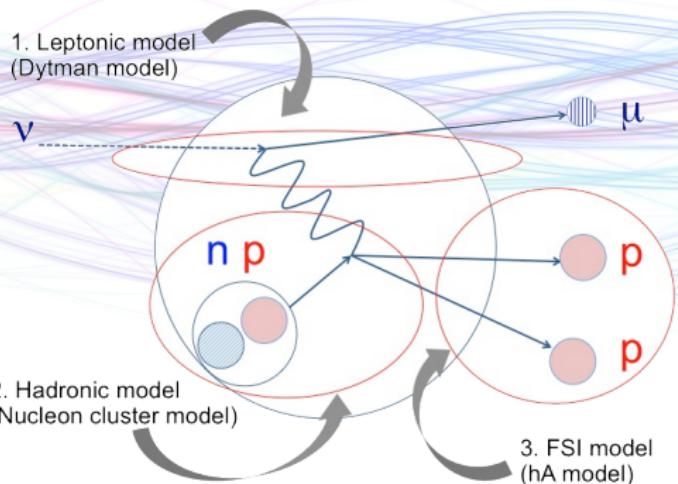
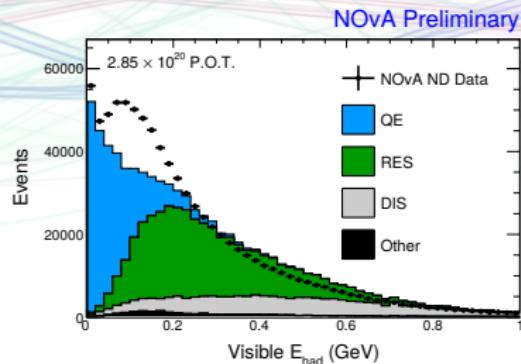
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- ▶ MINERvA report similar excess in their data¹

¹P.A. Rodrigues *et al.*, PRL 116 (2016) 071802 (arXiv:1511.05944)

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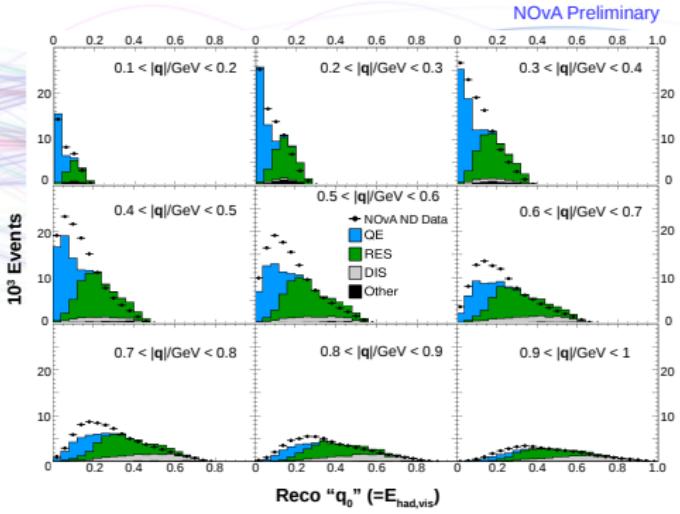
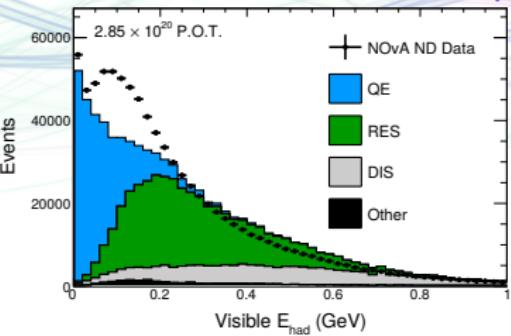
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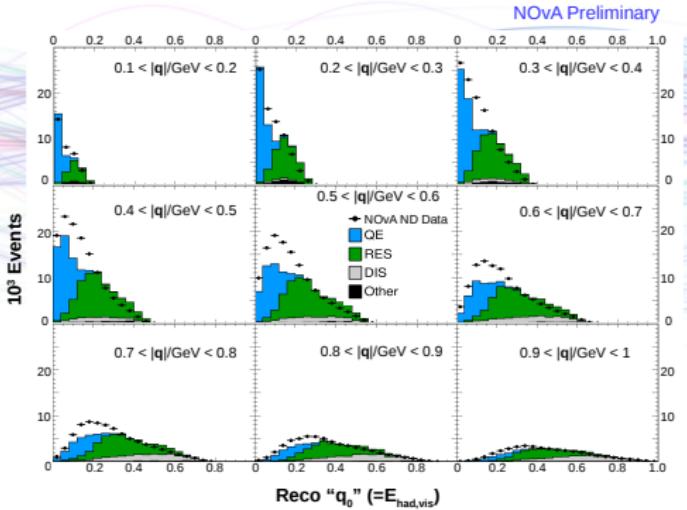
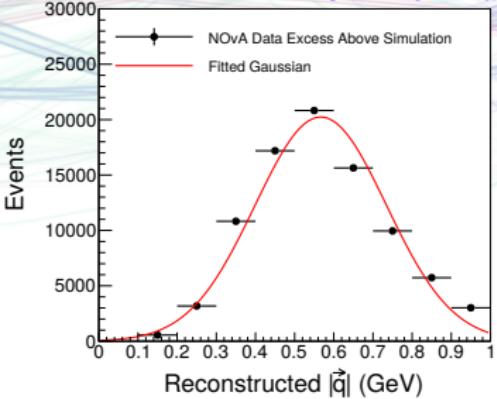
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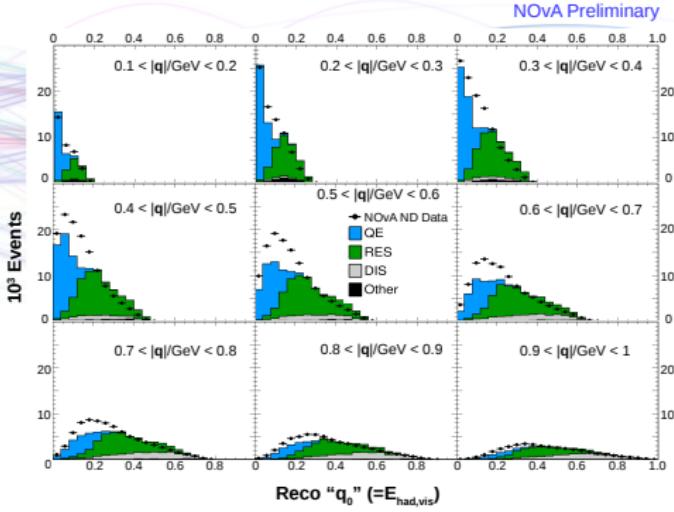
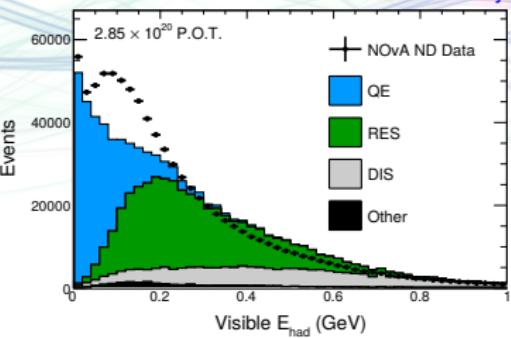
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- Enable GENIE's empirical **Meson Exchange Current** model²
- Reweighting to match observed excess as a function of \vec{p} transfer
- Also reduce single non-resonant pion production by 50%³

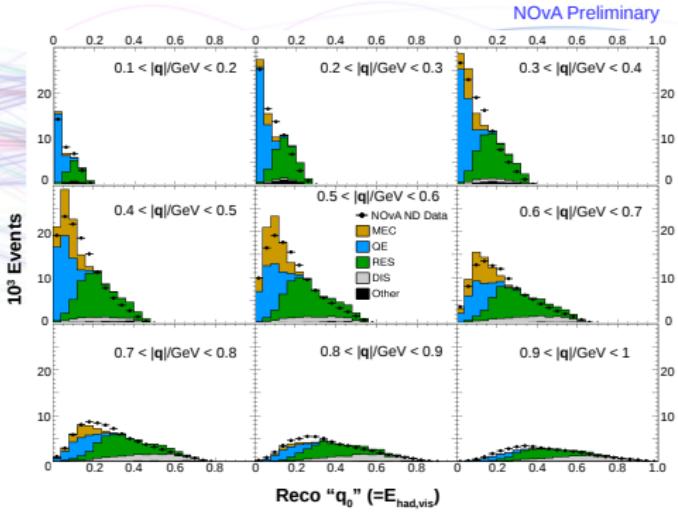
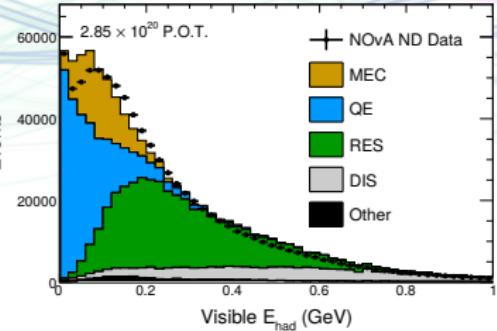
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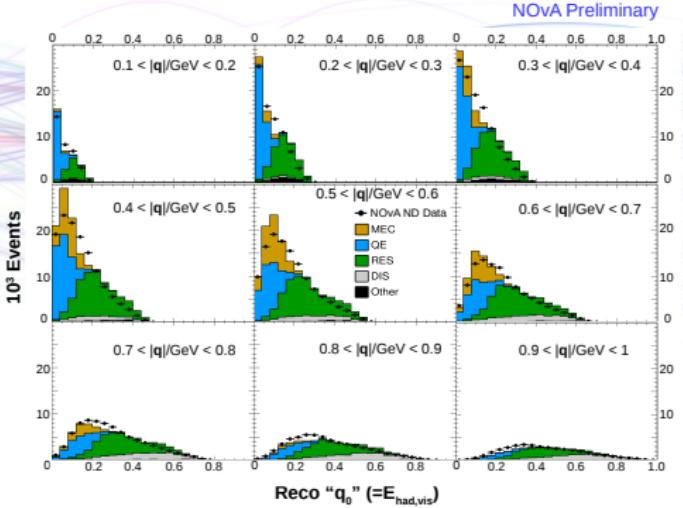
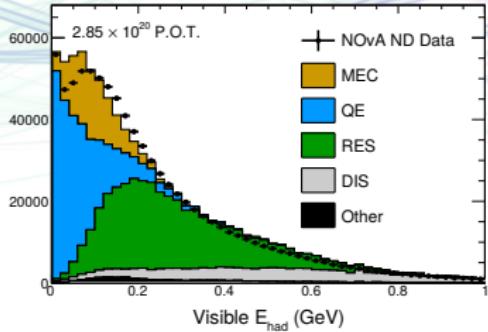
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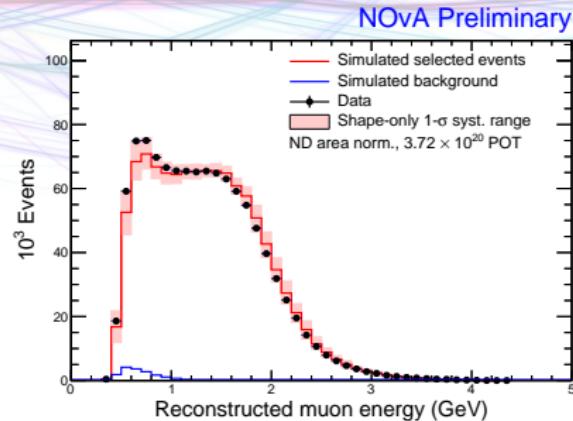
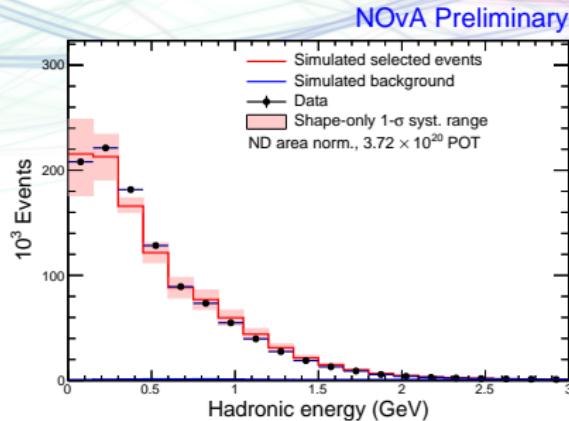
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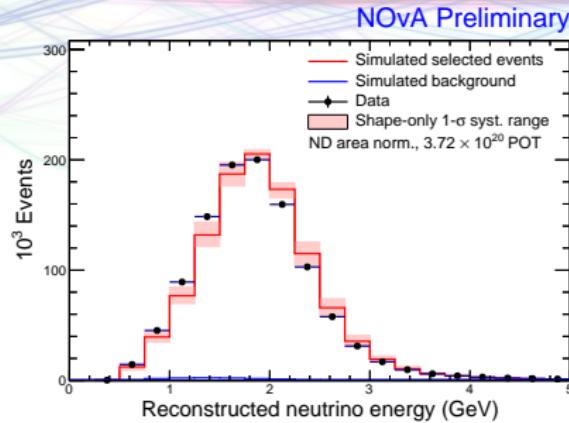
- ▶ Take 50% systematic on MEC component
- ▶ Reduced analysis' hadronic energy scale and QE xsec systematics

Muon neutrino energy reconstruction



- ▶ Good data/MC agreement for muon neutrino selected events
- ▶ Hadronic energy scale uncertainty improved to 5%

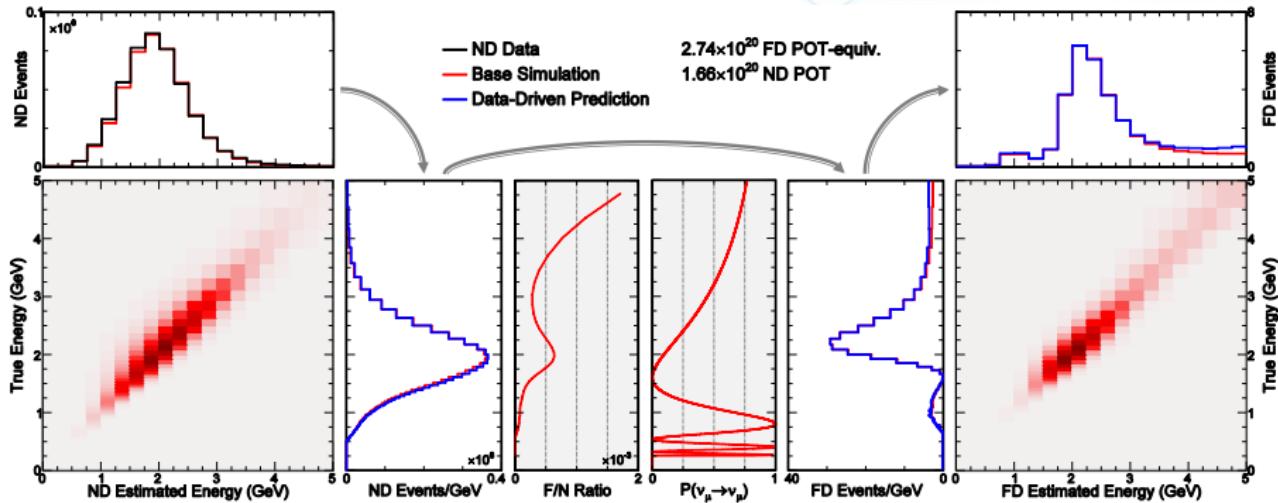
Muon neutrino energy reconstruction



- ▶ Good data/MC agreement for muon neutrino selected events
- ▶ Hadronic energy scale uncertainty improved to 5%
- ▶ Use ND data to predict FD neutrino spectrum

Extrapolation procedure

- ▶ Translate ND observations to true energy
- ▶ Transport to far detector and oscillate
- ▶ Smear back to reco energy



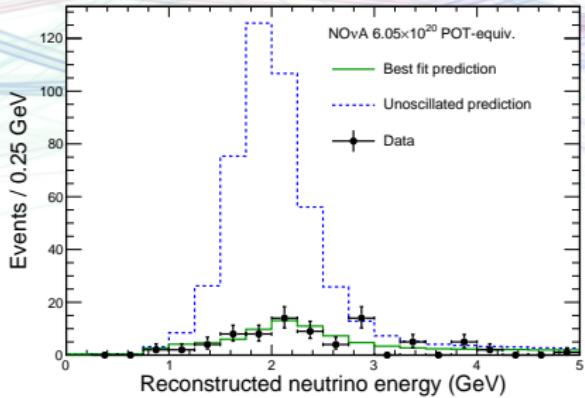
Systematic uncertainties

Source of uncertainty	Fractional uncertainty $\sin^2 \theta_{23}$ ($\pm\%$)	Fractional uncertainty Δm_{32}^2 ($\pm\%$)
Normalization	1.0	0.2
Muon E scale	2.2	0.8
Calibration	2.0	0.2
Relative E scale	2.0	0.9
Cross sections + FSI	0.6	0.5
Osc. parameters	0.7	1.5
Beam backgrounds	0.9	0.5
Scintillation model	0.7	0.1
Total systematic	3.4	2.4
Statistical uncertainty	4.1	3.5

- ▶ Consider multiple possible sources of systematic error
- ▶ Propagate effect of each through extrapolation
- ▶ Include as pull terms in fit
- ▶ Quoting increase (in quadrature) of measurement error

ν_μ disappearance results

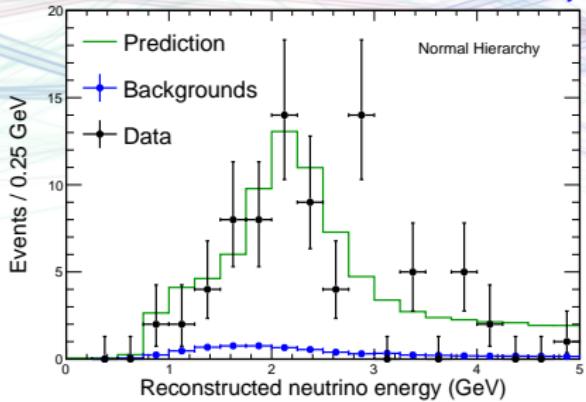
NOvA Preliminary



- ▶ Expect 473 FD ν_μ CC events with no oscillation
- ▶ Observe 82 (inc. 3.7 beam bkg. and 2.9 cosmic)

ν_μ disappearance results

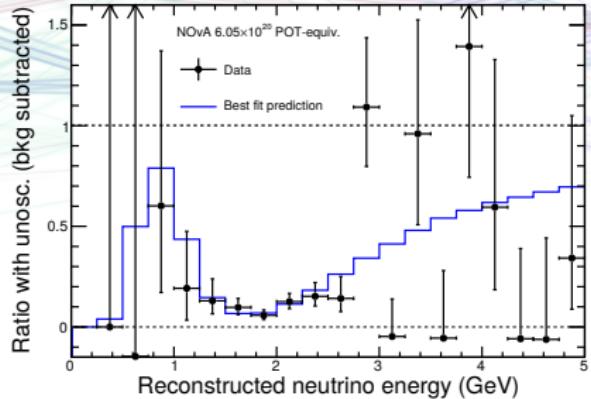
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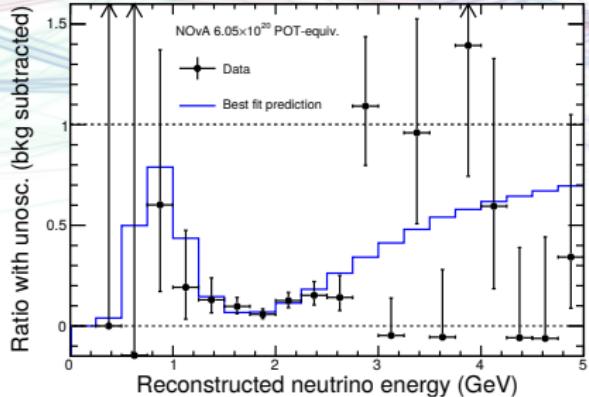
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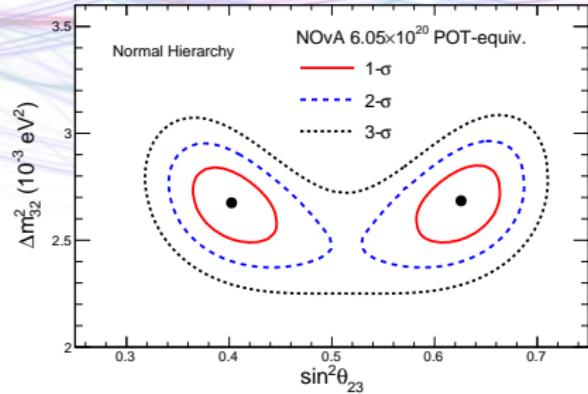
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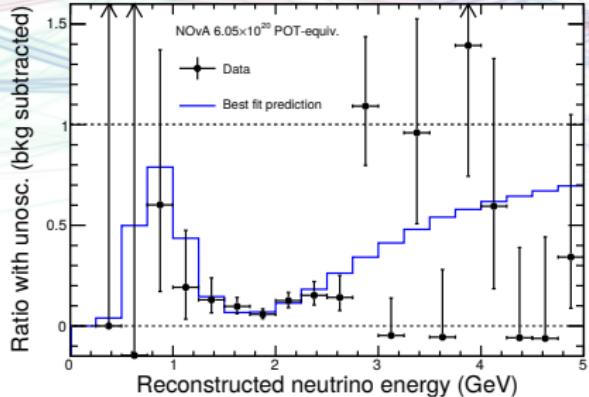
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$$\begin{aligned}\Delta m_{32}^2 &= (2.67 \pm 0.12) \times 10^{-3} \text{ eV}^2 \text{ (NH)} \\ \sin^2 \theta_{23} &= 0.40^{+0.03}_{-0.02} (0.63^{+0.02}_{-0.03})\end{aligned}$$

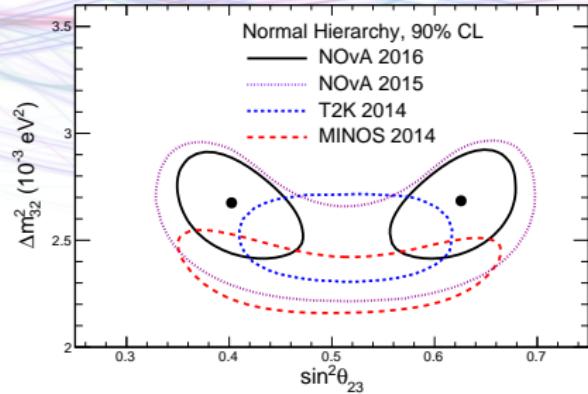
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ν_μ disappearance results

NOvA Preliminary



NOvA Preliminary



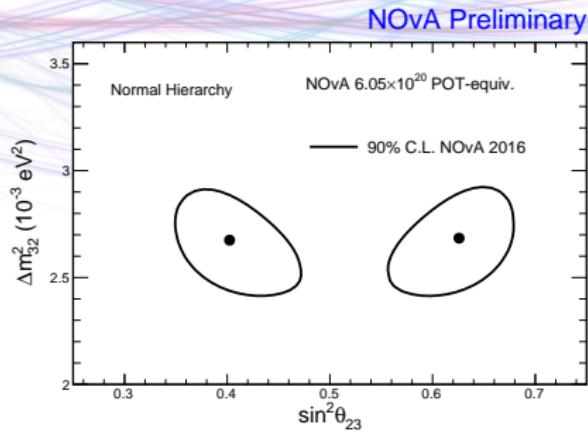
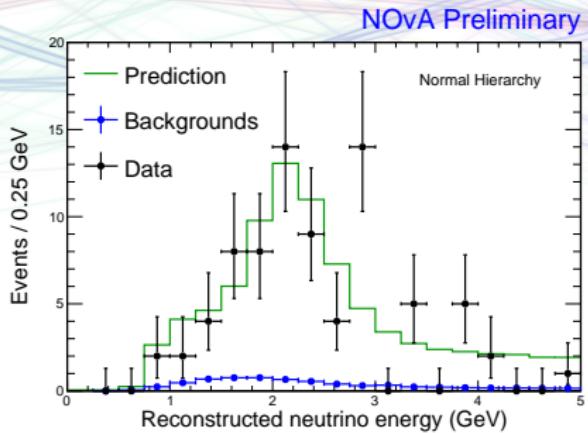
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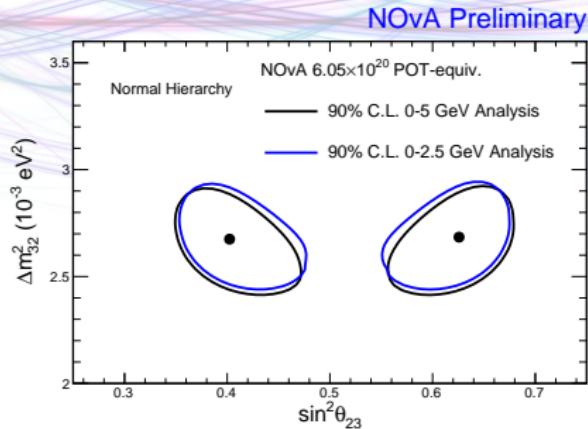
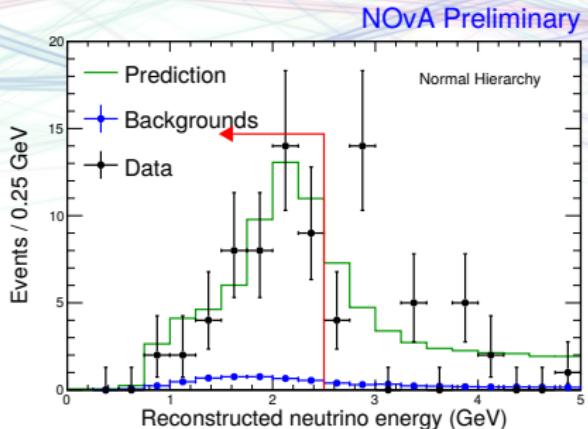
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ν_μ disappearance results



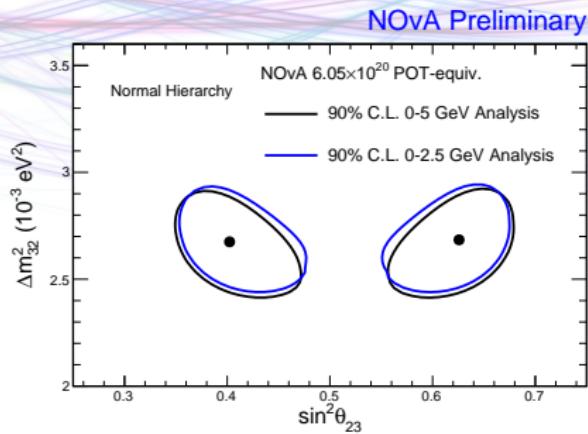
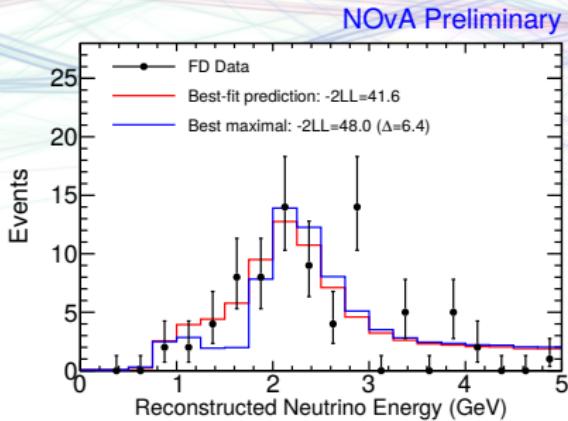
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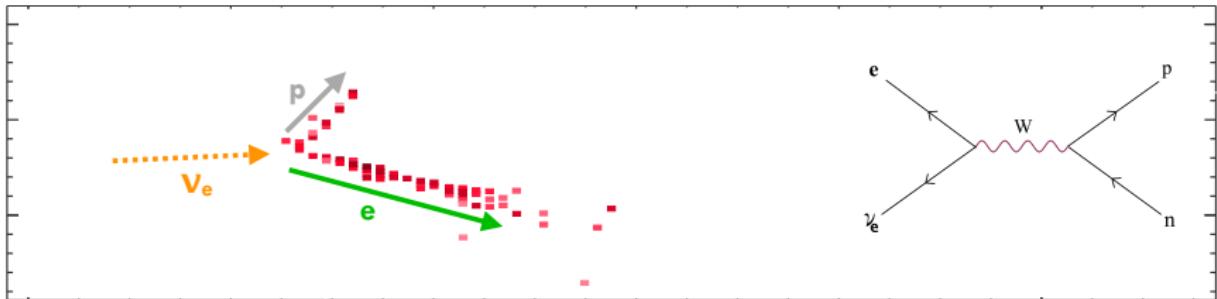


- $\chi^2/\text{dof} = 41.5/17$, driven by fluctuations in the tail
- No significant pull on oscillation fit from these bins
- Non-maximal best fit driven by bins in oscillation dip, as expected
 - $\Delta m_{32}^2 = 2.46 \times 10^{-3} \text{ eV}^2$ when forcing maximal mixing
 - $\Delta\chi^2 = 6.4$ above non-maximal fit

Principle of the ν_e measurement



- ▶ Separate ν_e CC interactions from beam backgrounds
 - ▶ Harder problem than ν_μ CC selection
- ▶ Evaluate remaining backgrounds in ND
 - ▶ Intrinsic beam ν_e
 - ▶ Neutral currents
 - ▶ ν_μ CC – mostly oscillates away
- ▶ An excess in the FD is the sign of $\nu_\mu \rightarrow \nu_e$ oscillations



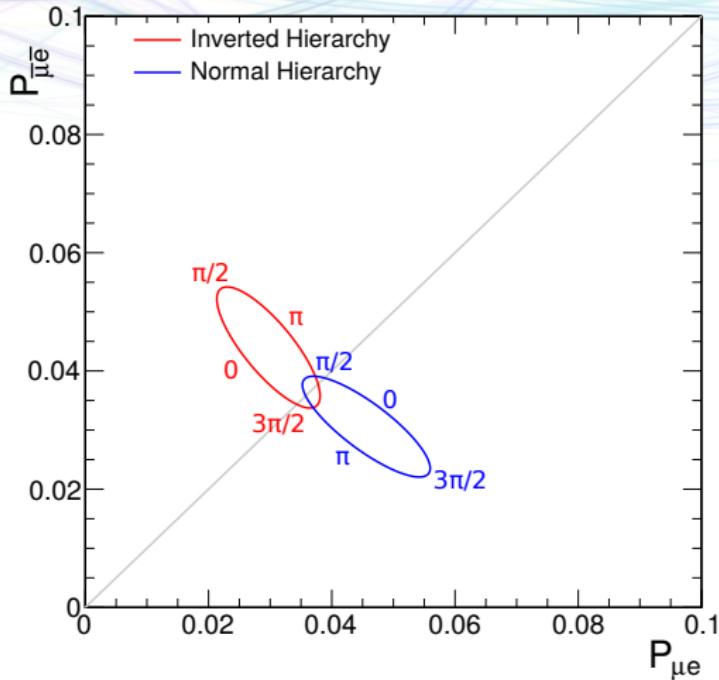
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- ▶ An excess in the FD is the sign of $\nu_\mu \rightarrow \nu_e$ oscillations
- ▶ $P_{\mu e} \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E} \right) + f(\text{sign}(\Delta m_{32}^2)) + f(\delta_{CP})$
- ▶ θ_{13} only 8.5° degrees, most ν_μ go to ν_τ instead
- ▶ Look for deviations due to hierarchy (matter effects) and CP-violation

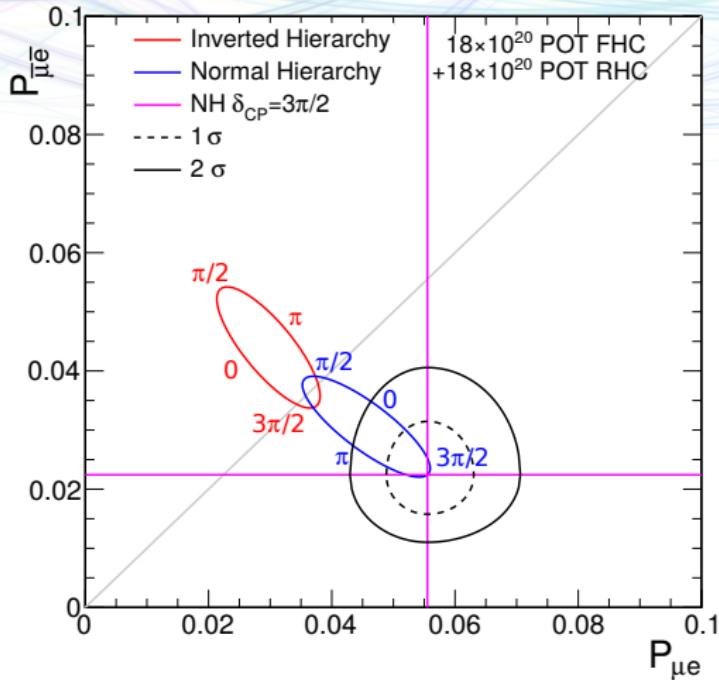
Principle of the ν_e measurement

- ▶ To first order, NOvA measures $P(\nu_\mu \rightarrow \nu_e)$ and $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$ evaluated at 2GeV
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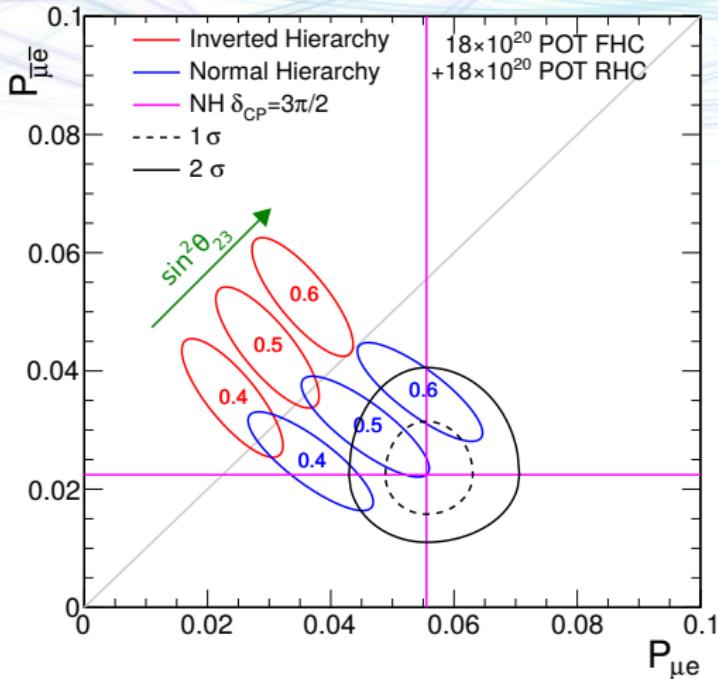
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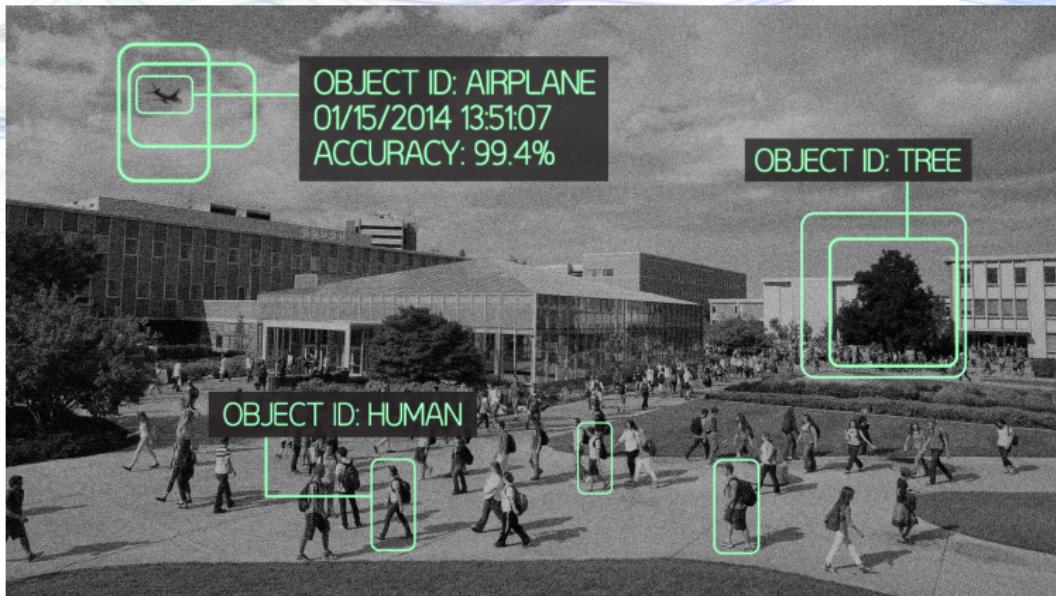


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- ▶ P also $\propto \sin^2 \theta_{23}$



Convolution Neural Networks



- ▶ Recent advances in machine learning/computer vision
- ▶ Achieving near-human performance on image classification tasks
- ▶ Why not classify event-displays?

Convolution Neural Networks

$$\frac{1}{8} \begin{bmatrix} -1 & -1 & -1 \\ -1 & +8 & -1 \\ -1 & -1 & -1 \end{bmatrix}$$

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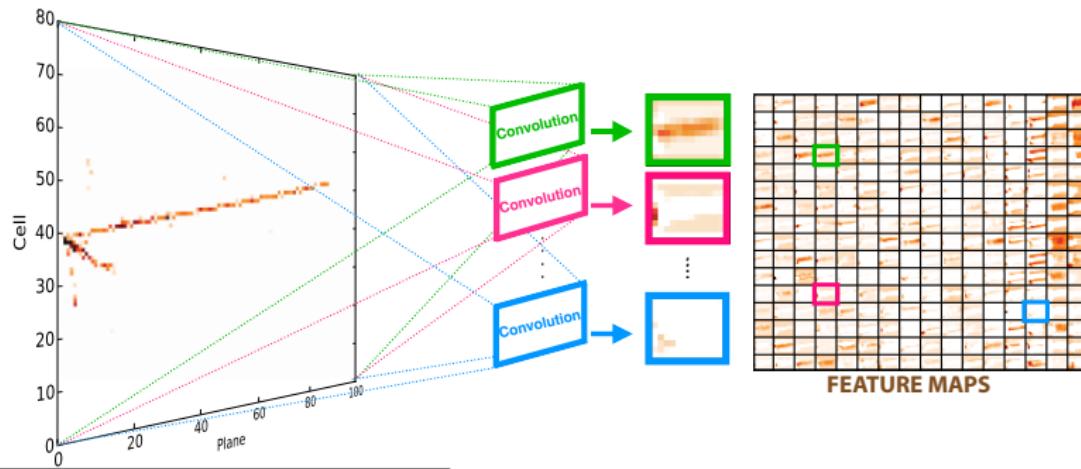
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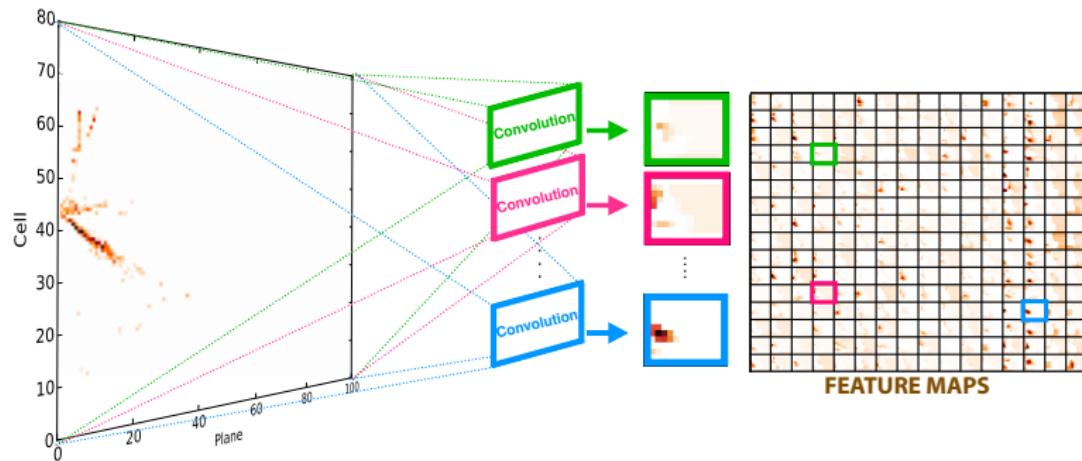
Selecting electron neutrinos – CVN

- ▶ Convolutional Visual Network (CVN)
- ▶ Early layers perform convolutions to pick out abstract features
- ▶ Fully-connected final layers
- ▶ Trained using FNAL's Wilson Cluster GPUs
- ▶ Statistical power equivalent to 30% more exposure than previous IDs



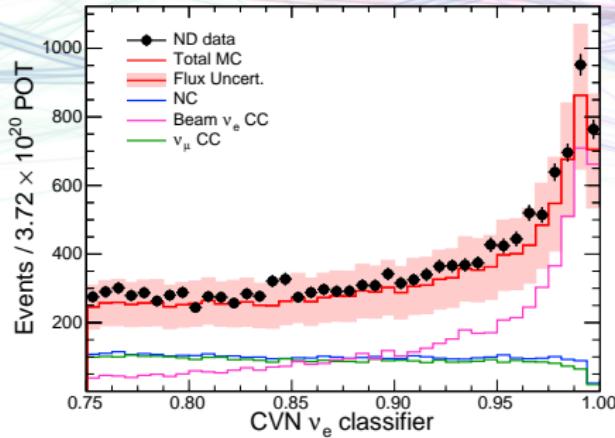
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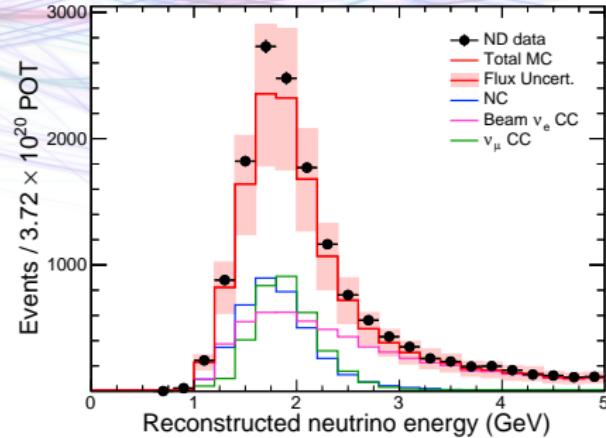


Selecting electron neutrinos – CVN

NOvA Preliminary



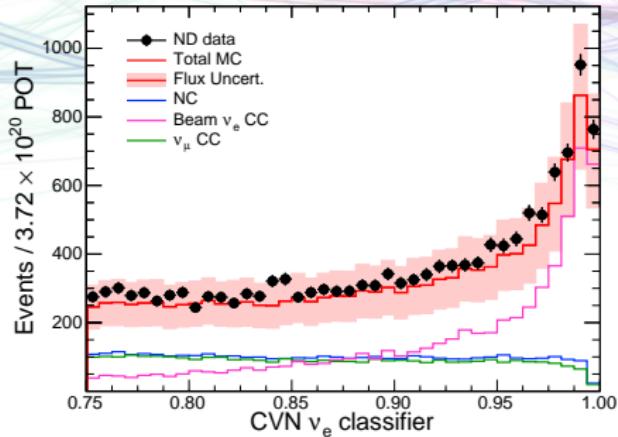
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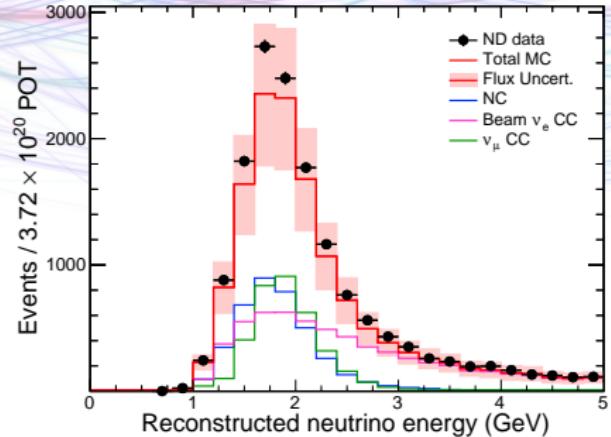
- ▶ 73% ν_e signal efficiency, 76% purity
- ▶ Loosen PID cut to maximize $s/\sqrt{s+b}$
- ▶ Analyze in 3 PID \times 4 energy bins

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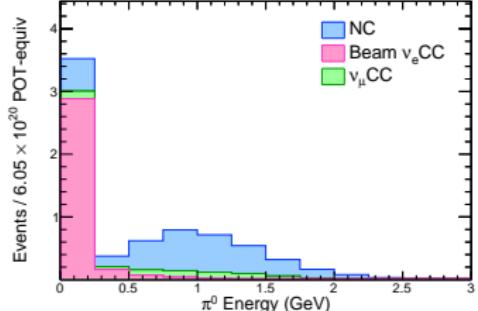
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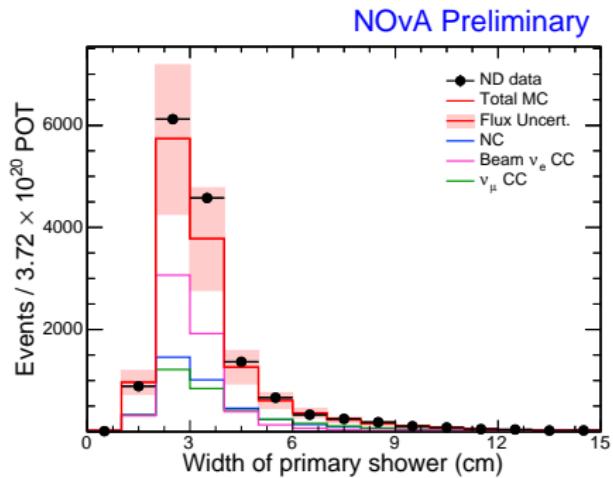
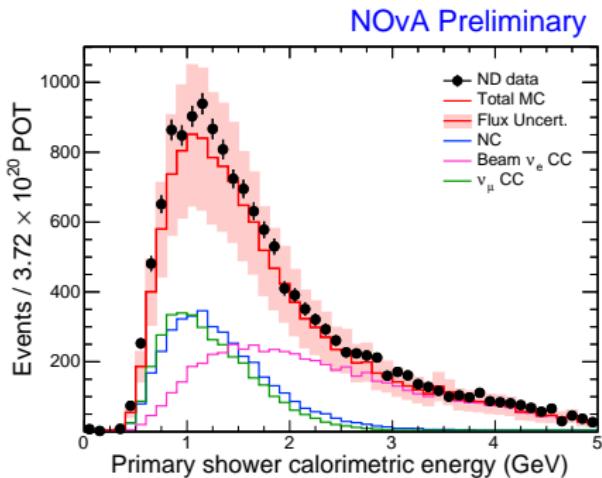
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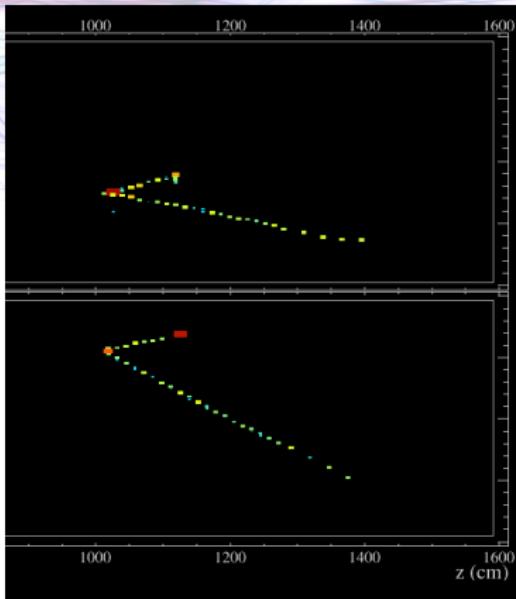
Selector checks

- ▶ CVN subjected to the same (or greater) scrutiny than past PIDs
- ▶ ND data/MC is good
- ▶ Better cosmic rejection and similar systematics to other PID options



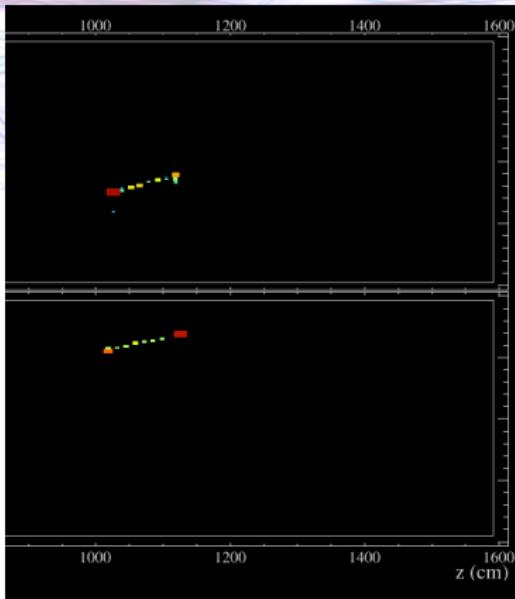
ν_e selection efficiency – MRE

- ▶ EM showers should be well modelled
- ▶ Any ν_e signal efficiency differences coming from the hadronic side?
- ▶ Remove muon from clear ν_μ CC events in ND, replace with simulated shower
- ▶ $\mathcal{O}(1\%)$ efficiency difference to select MRE data/MC events



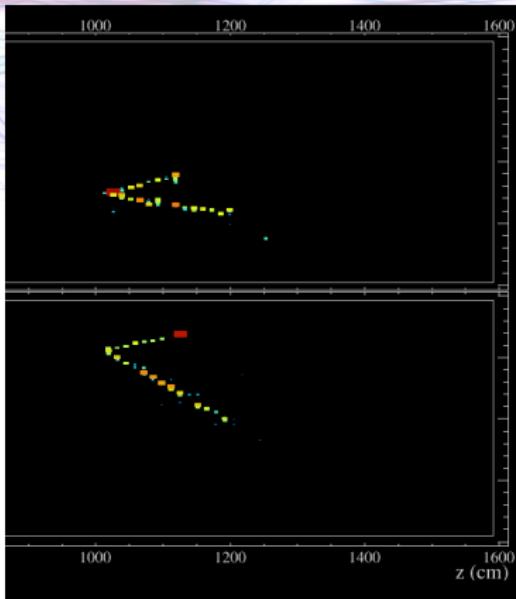
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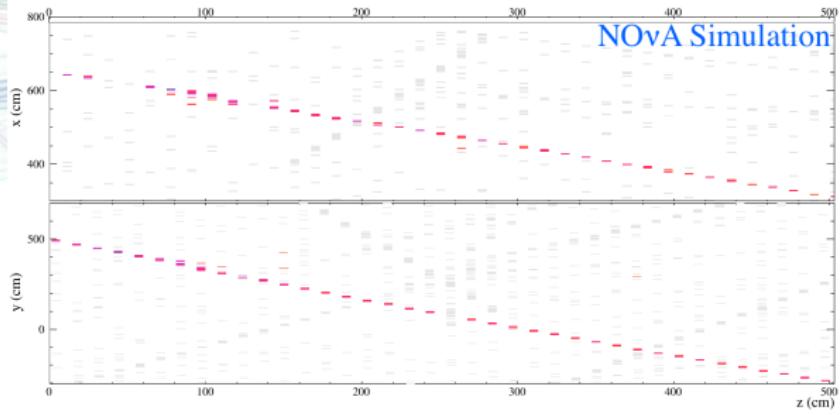


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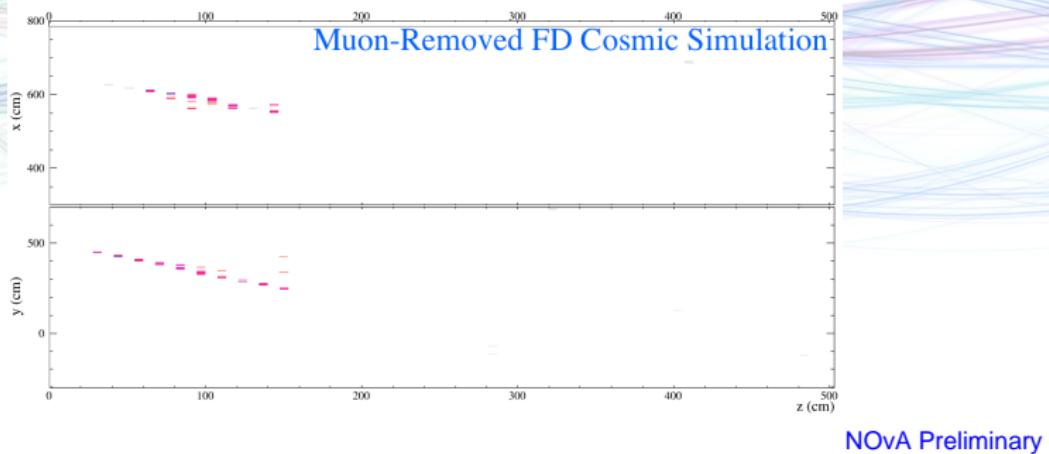


ν_e selection efficiency – EM activity

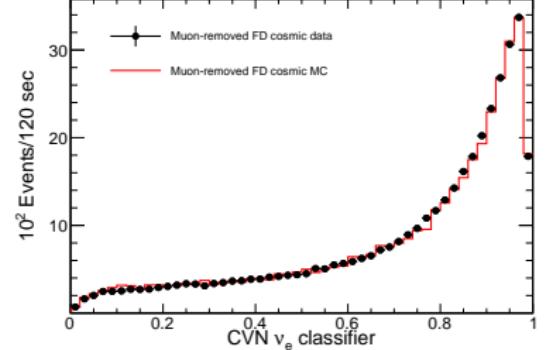


- ▶ Find FD data cosmic rays w/ brems

ν_e selection efficiency – EM activity



- ▶ Find FD data cosmic rays w/ brems
- ▶ Remove μ leaving pure EM activity
- ▶ Run through PID in data and MC
- ▶ Very good agreement

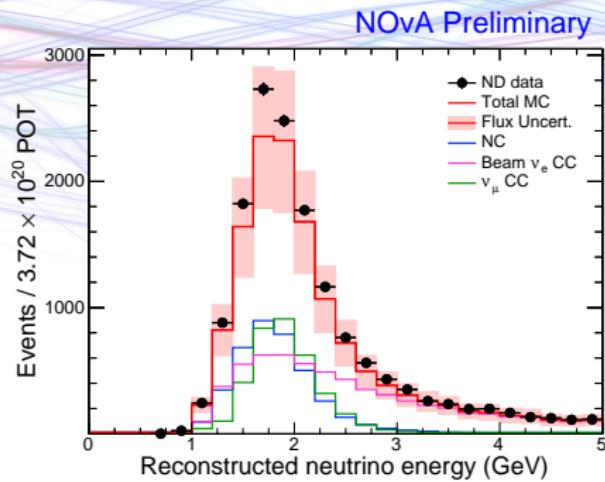


ND decomposition

- ▶ Use ND data to predict FD backgrounds
 - ▶ Beam ν_e CC
 - ▶ NC
 - ▶ ν_μ CC
- ▶ $\sim 10\%$ excess of data over MC

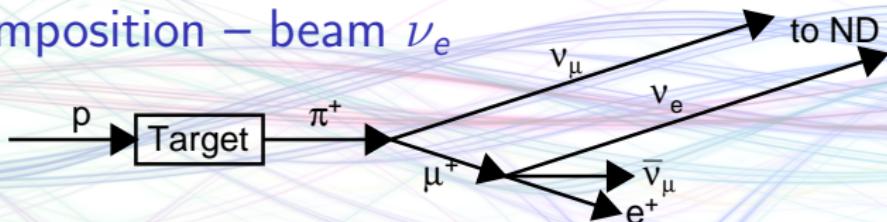


C. Backhouse (Caltech)

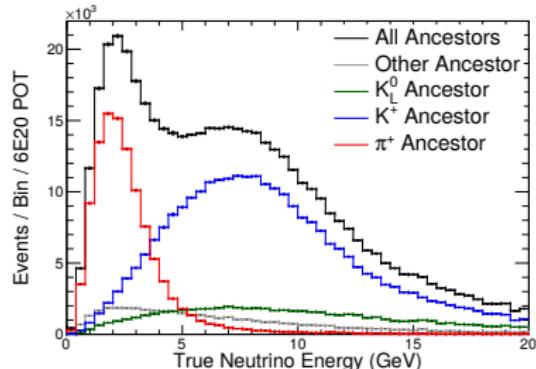
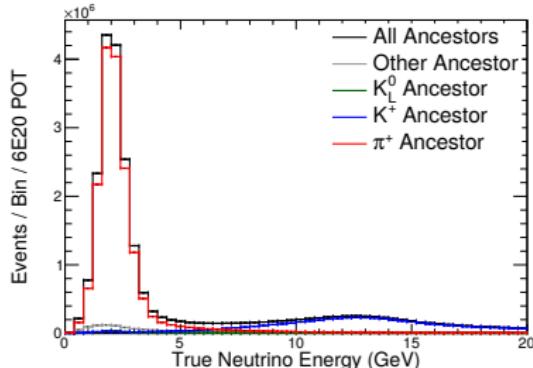


- ▶ How to divide between the components?
- ▶ e.g. most ν_μ CC oscillate away before FD

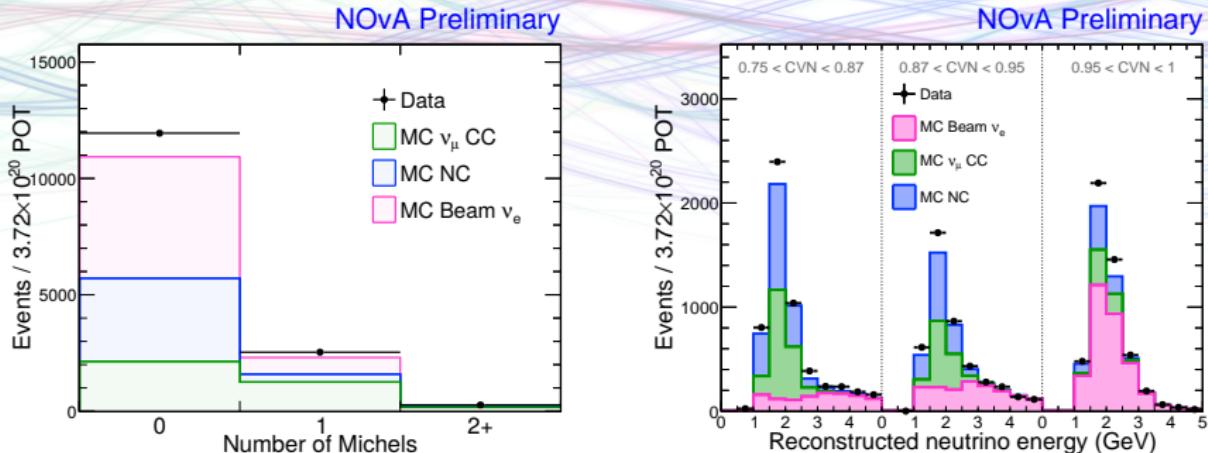
ND decomposition – beam ν_e



- ▶ Low- E ν_μ and ν_e trace back to the same π^+ ancestors
- ▶ Weight ν_e with K^+ parents up 17% based on ν_μ high- E tail
- ▶ Use ν_μ at lower energy to reweight decaying pions in (p_T, p_z) space
- ▶ Decreases ν_e with π^+ parent 3-4%
- ▶ Overall effect is 3% increase in 1-3 GeV range

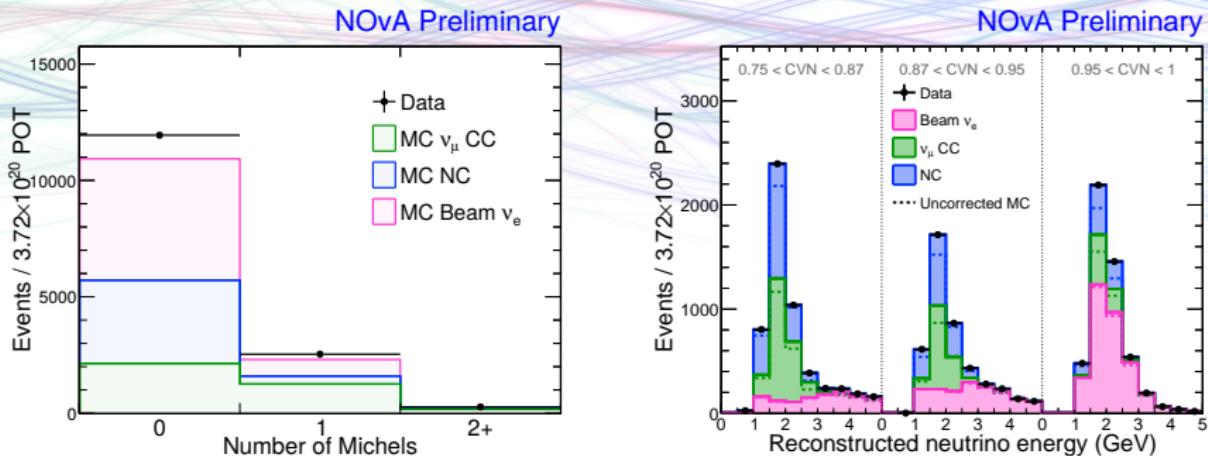


ND decomposition – Michels



- ▶ ν_μ CC background events have Michel electron from muon decay
- ▶ Also produced in ν_e CC and NC by pions, but ν_μ have ~ 1 more
- ▶ Fit observed N_{michel} spectrum in each bin by varying components
- ▶ ν_e and NC near-degenerate, fix ν_e to parent-reweight estimate
- ▶ Data excess assigned between NC (+17%) and ν_μ CC (+10%)

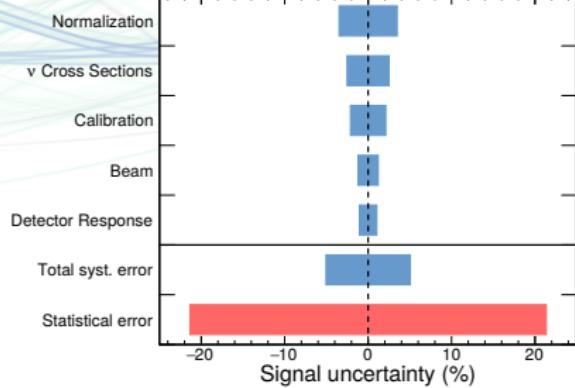
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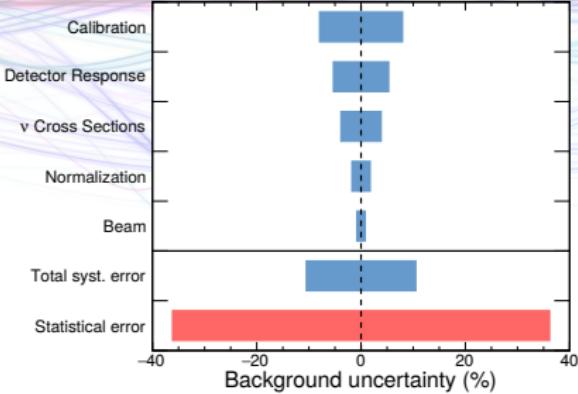
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Systematics

Signal



Background



- ▶ Considered multiple possible sources of systematic error
- ▶ Propagate shifts through to update FD prediction
- ▶ Total $\sim 5\%$ error on signal, 10% on bkg.
- ▶ Fit nuisance parameters as pull terms
- ▶ Dominated by statistical error

Event count expectations

Total prediction

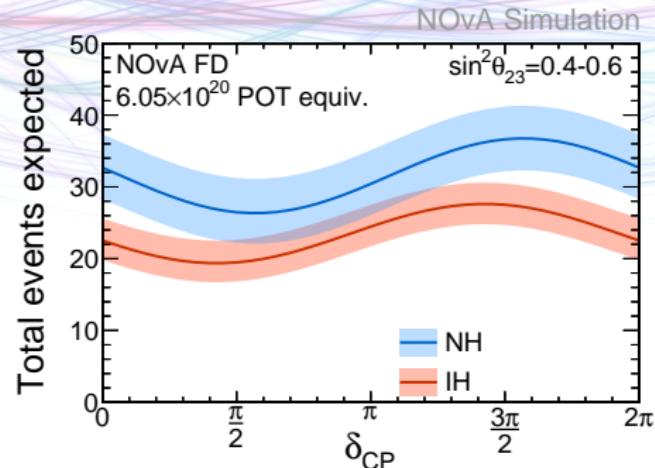
$$\begin{array}{ll} \text{NH } \frac{3\pi}{2} & \text{IH } \frac{\pi}{2} \\ 36.4 & 19.4 \end{array} \quad \sin^2 \theta_{23} = 0.5 \quad \pm 5\% \text{ syst.}$$

$$P_{\nu_\mu \rightarrow \nu_e} \propto \sin^2 \theta_{23} \sin^2 2\theta_{13}$$

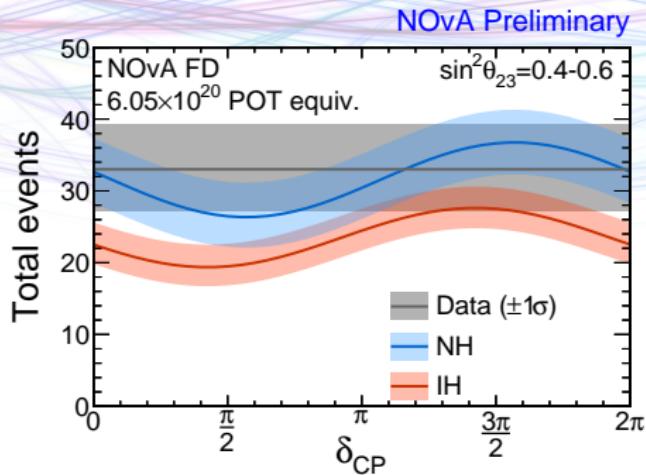
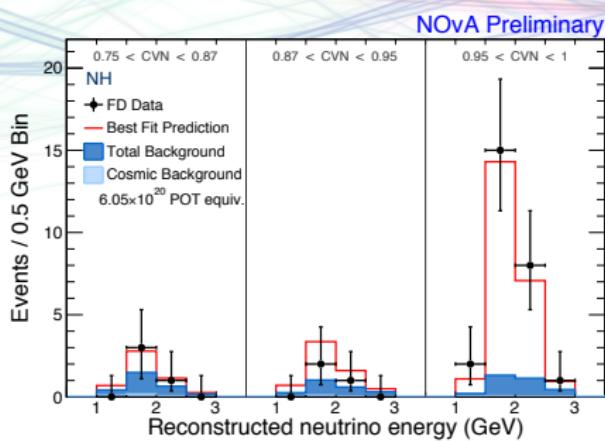
Background components

Total bkg	NC	beam	ν_e	ν_μ	CC	ν_τ	CC	cosmics	$\pm 10\%$ syst.
8.2	3.7		3.1	0.7		0.1	0.5		

Essentially independent of oscillation parameters



ν_e appearance results

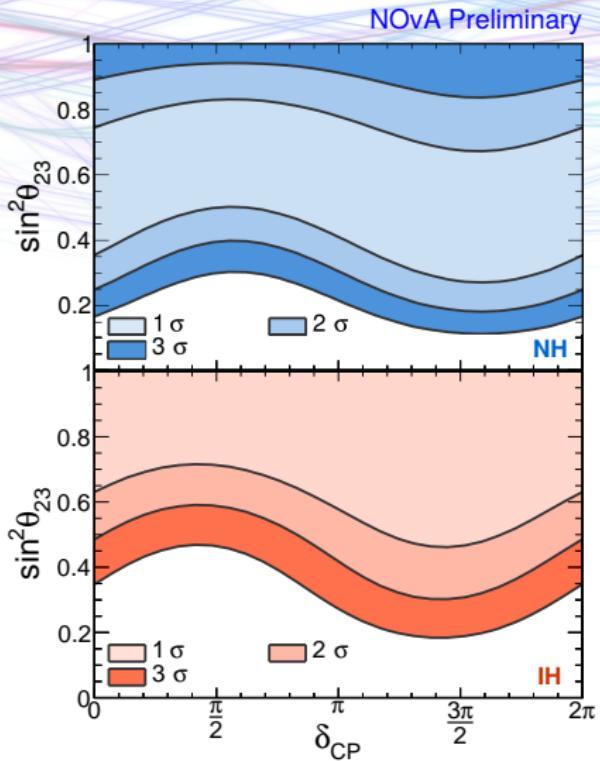


- ▶ Observe **33** events passing ν_e selection
- ▶ On 8.2 background
- ▶ Towards the higher end of expectations

Previous result PIDs: LID(LEM) sees 34(33) events on bkg. of 12.2(10.3)

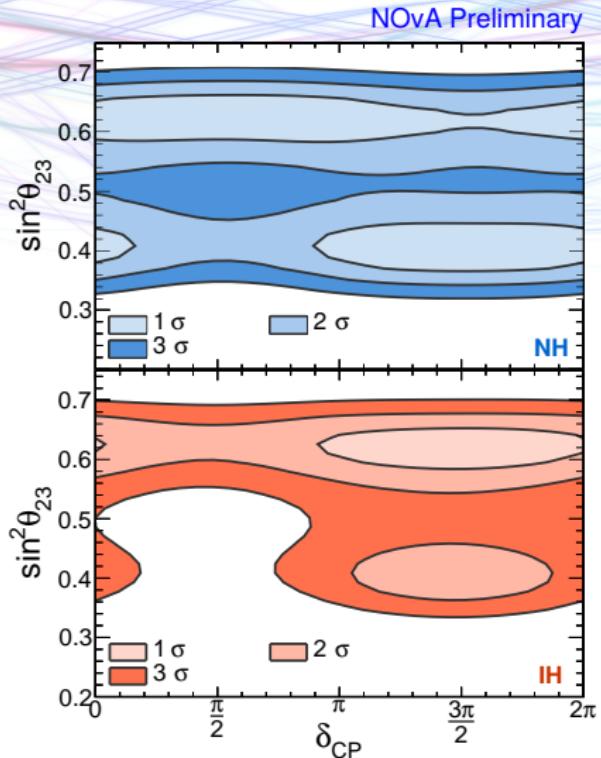
ν_e fit results

- ▶ Constrain θ_{13} to reactor average
 $\sin^2 2\theta_{13} = 0.085 \pm 0.005$



ν_e fit results

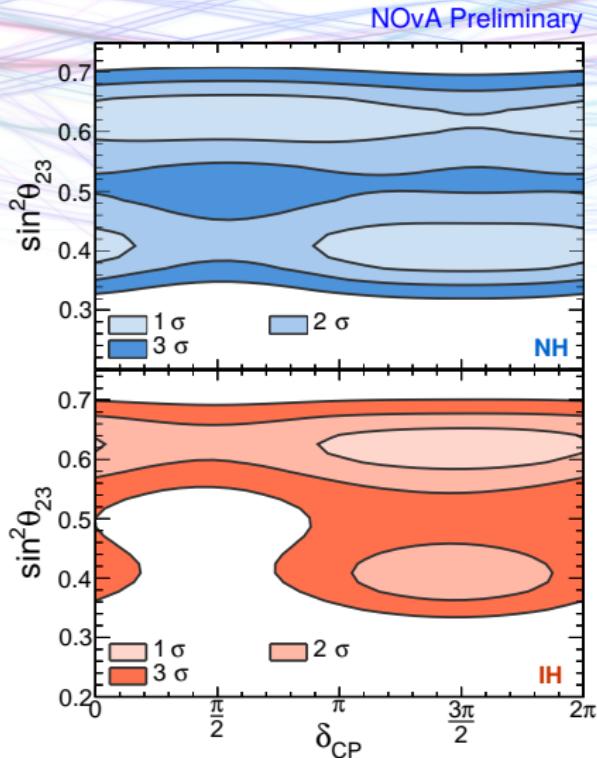
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- ▶ Add $\Delta m_{32}^2/\theta_{23}$ results from ν_μ analysis
- ▶ Not a full joint fit. No syst./osc. param correlations. No FC.



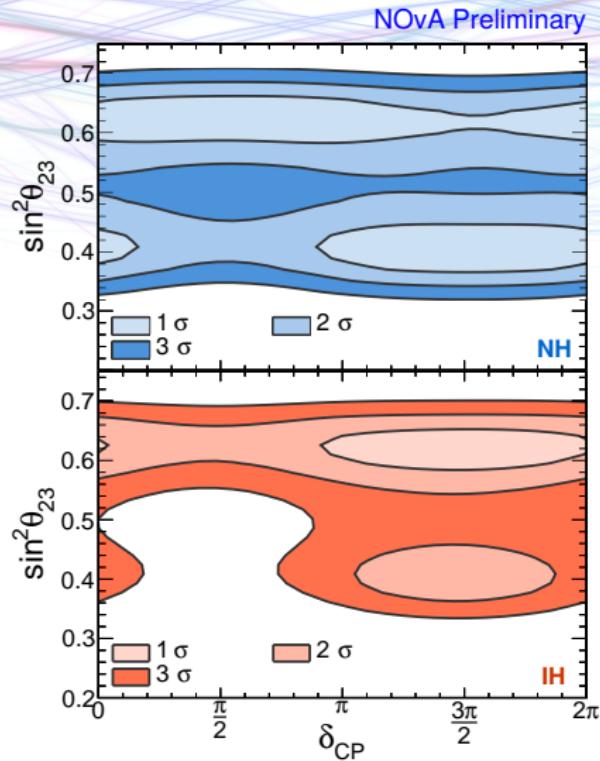
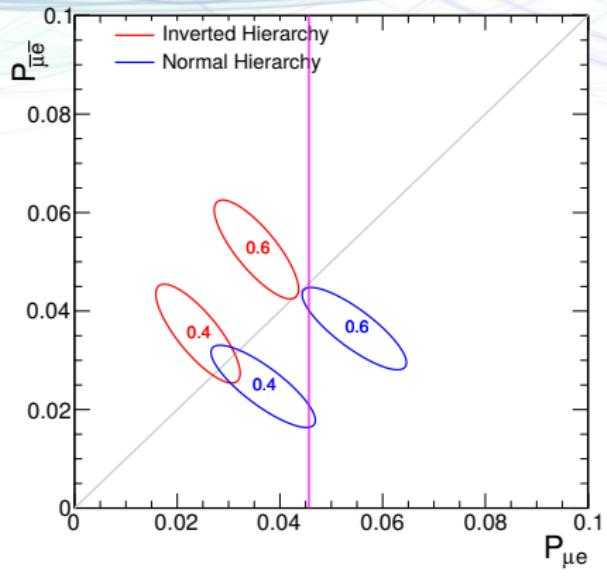
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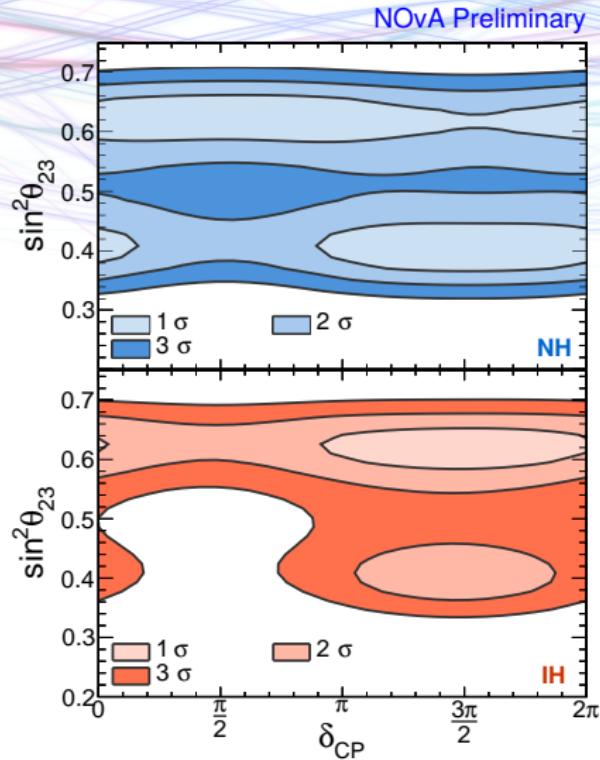
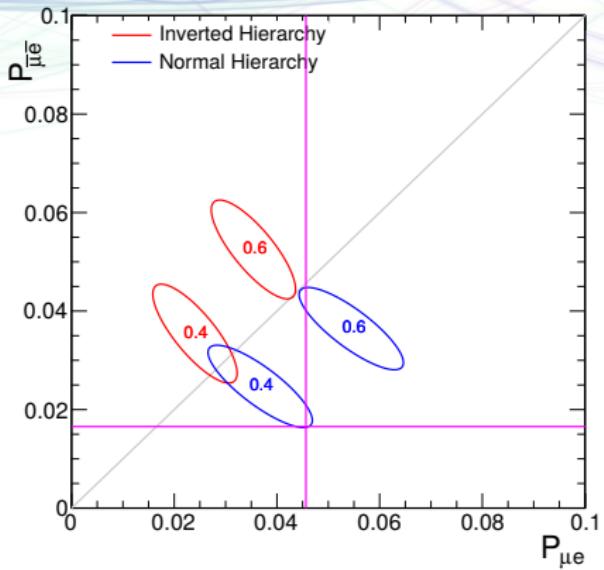
- ▶ Prefer NH, not statistically significant $\Delta\chi^2 = 0.46$
- ▶ Exclude region in IH, lower octant, around $\delta_{CP} = \pi/2$ at 3σ



ν_e fit results

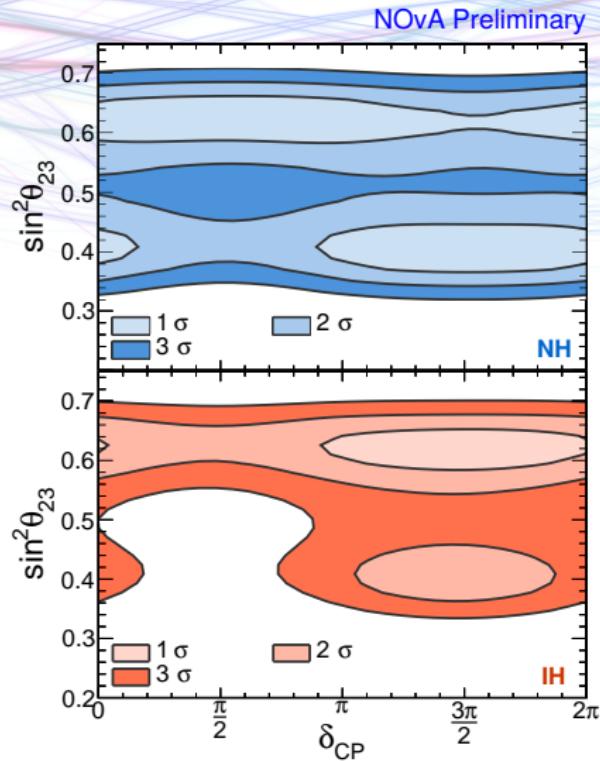
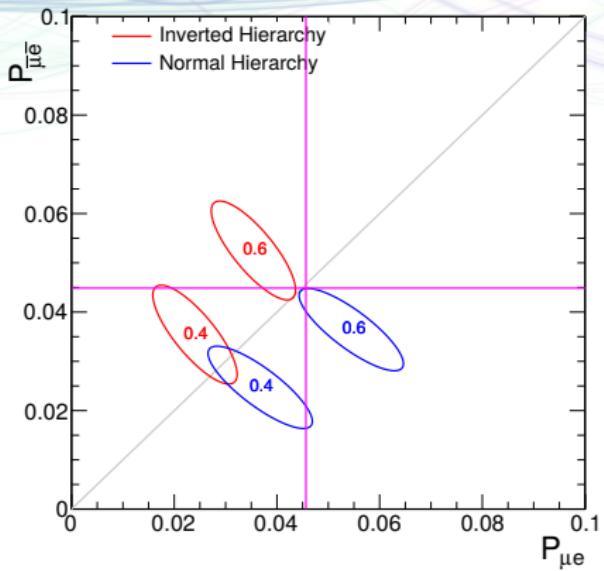


ν_e fit results



- Antineutrino data will help to resolve degeneracies
- $> 2\times$ difference in $\bar{\nu}_e$ rate between solutions

ν_e fit results



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Conclusion

- ▶ Muon neutrinos disappear (for sure)
- ▶ But not quite maximally (2.5σ)
- ▶ Electron neutrinos appear ($> 8\sigma$)
- ▶ Rate towards upper end of expected range
- ▶ Slight preference for NH, octant degenerate
- ▶ IH/lower/ $\frac{\pi}{2}$ ruled out ($> 3\sigma$)
- ▶ Will run antineutrinos in Spring 2017
- ▶ Thanks to everyone at Fermilab who contributed to NOvA's success!



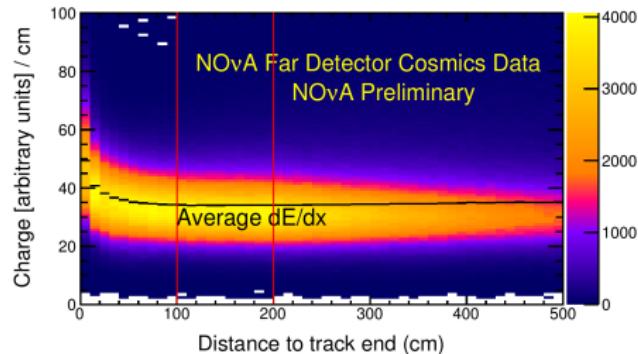
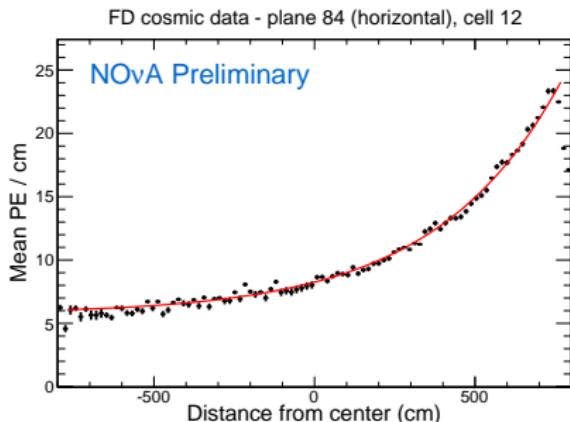
www-nova.fnal.gov

Thank you!

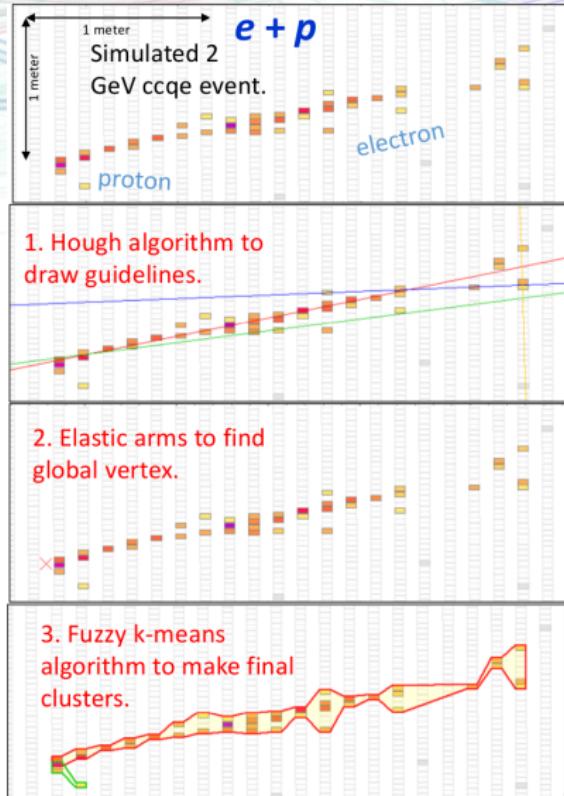
Questions?

Calibration and energy scale

- ▶ Detector response varies substantially over length of cell due to attenuation in fiber
- ▶ Use cosmic ray muons as a standard candle
- ▶ Calibrate every channel (344,064) individually
- ▶ Use dE/dx near the end of stopping muons to set absolute scale

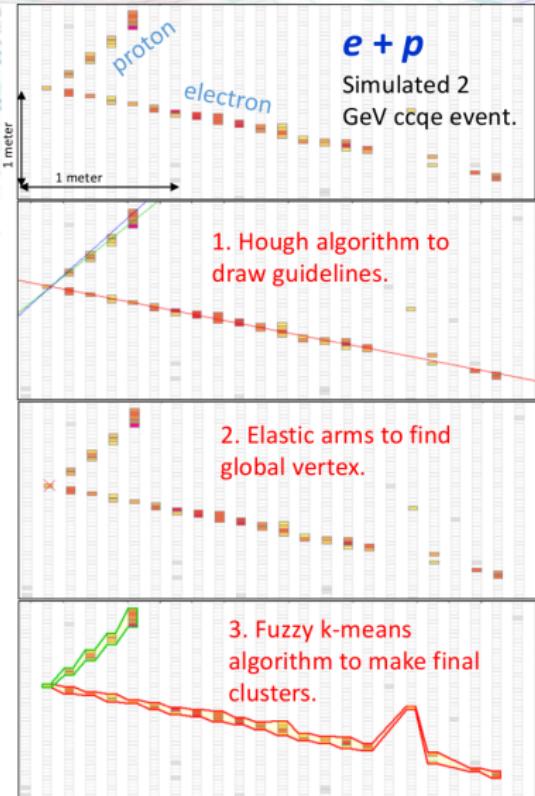


Event reconstruction



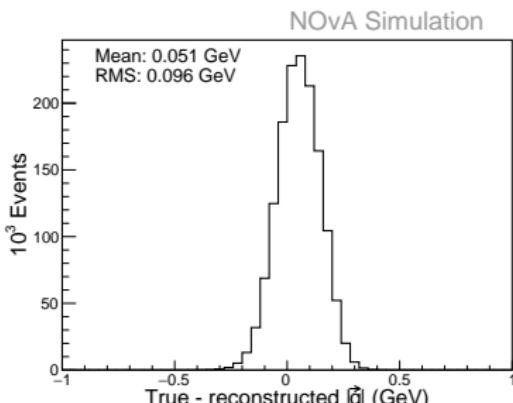
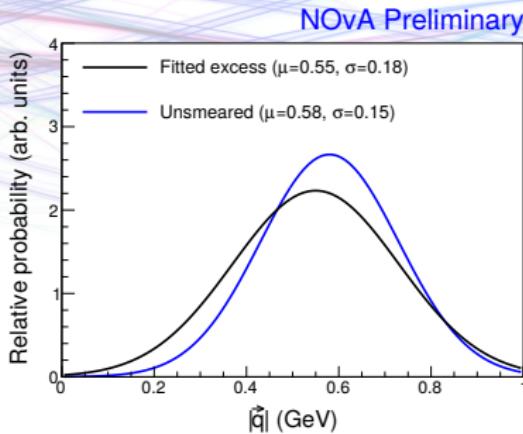
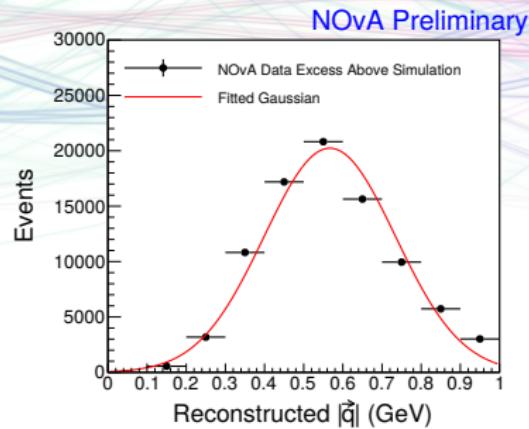
- ▶ First cluster hits in space and time
- ▶ Start with 2-point Hough transform
 - ▶ Line-crossing are vertex seeds
- ▶ ElasticArms finds vertex
- ▶ Fuzzy k -means clustering forms prongs
- ▶ ν_μ analysis uses a Kalman filter to reconstruct any muon track

Event reconstruction



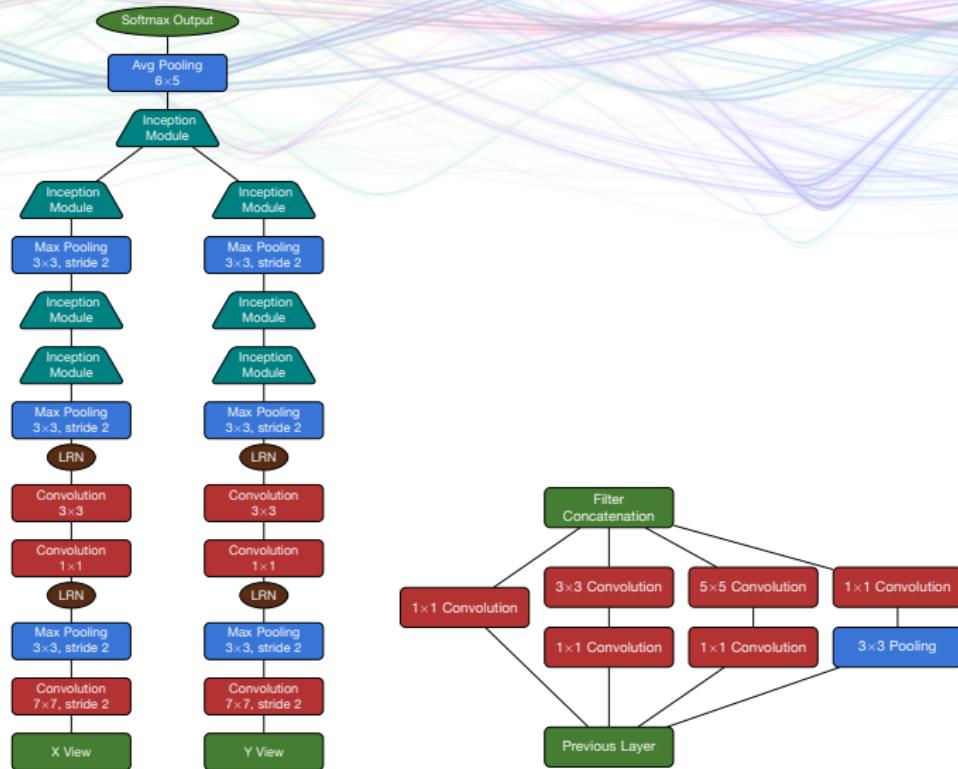
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MEC details

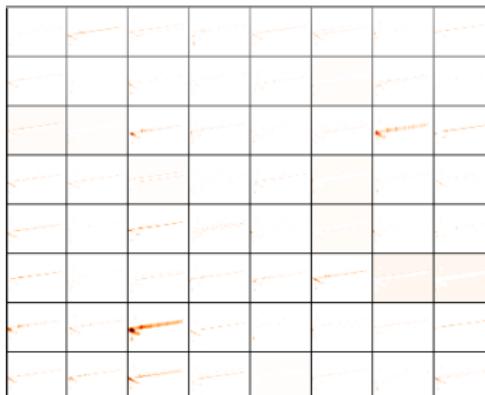
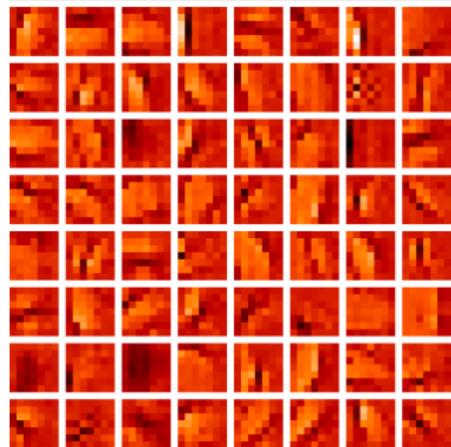
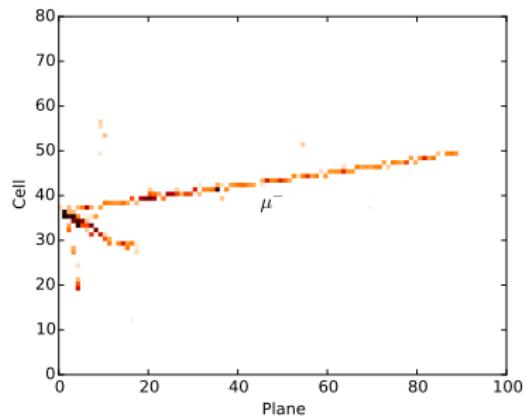


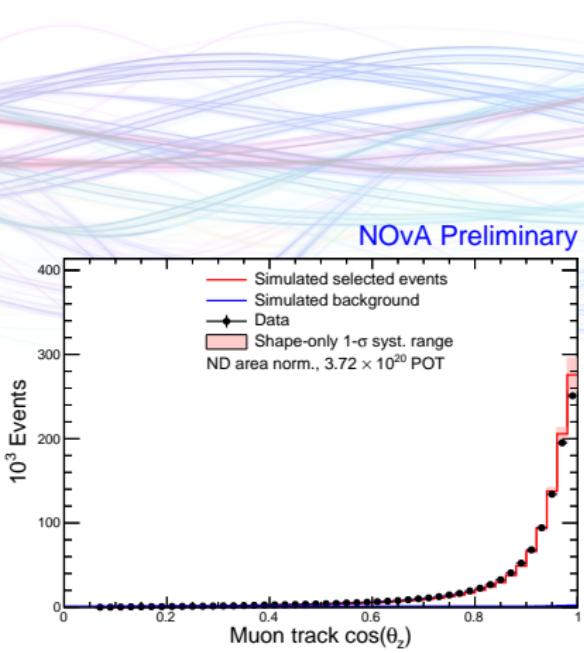
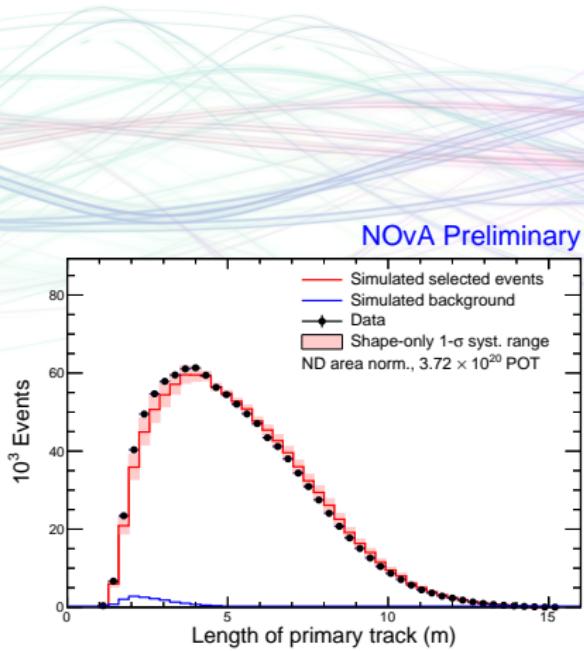
- ▶ Match q^0 distribution to GENIE's QE distribution
- ▶ Take RPA (nuclear charge screening) on/off as a systematic

CVN architecture



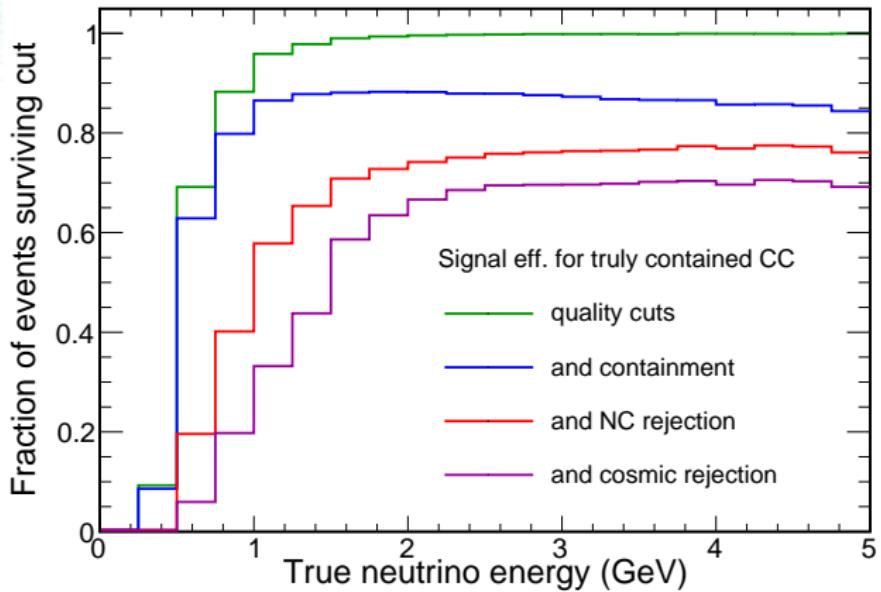
CVN example





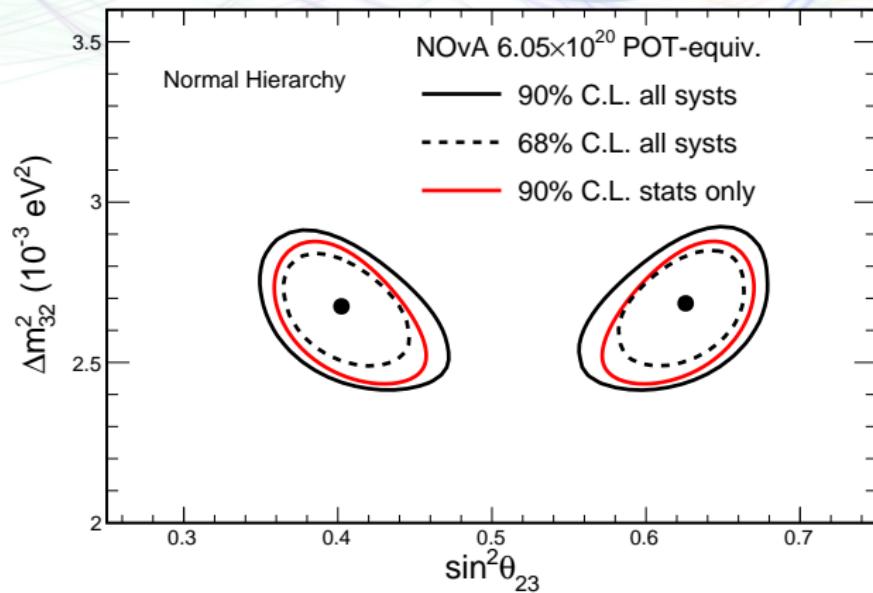
ν_μ selection efficiency

NOvA Preliminary



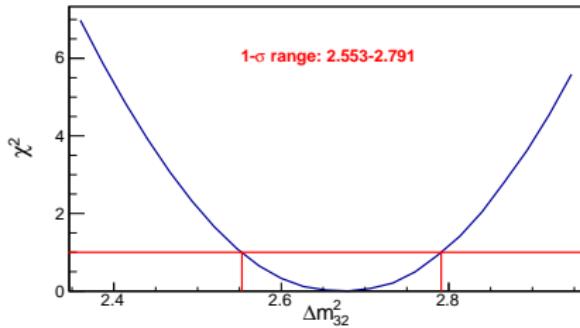
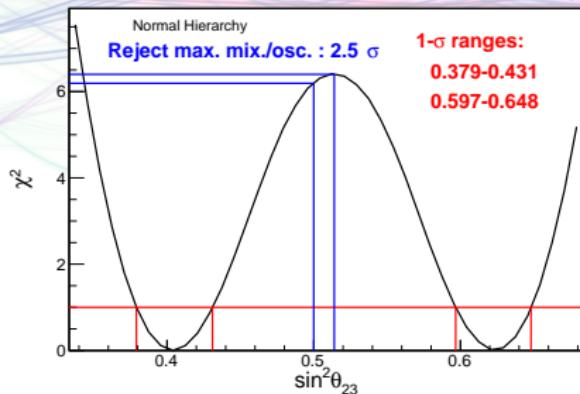
ν_μ impact of systematics

NOvA Preliminary



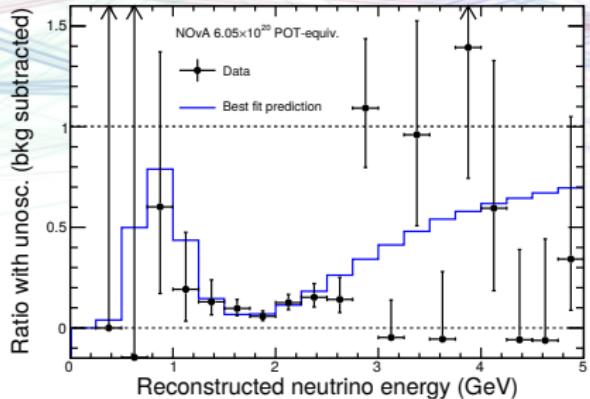
ν_μ 1D profiles

NOvA Preliminary

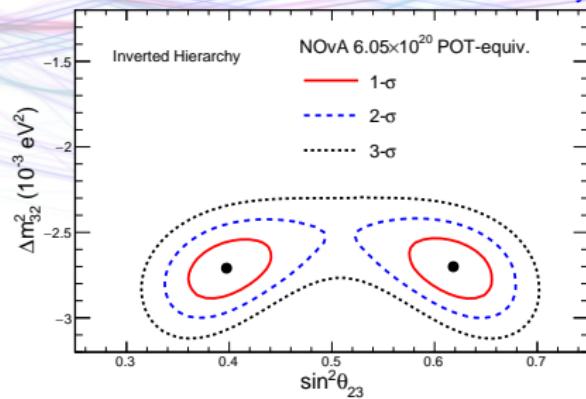


ν_μ disappearance results – IH

NOvA Preliminary



NOvA Preliminary



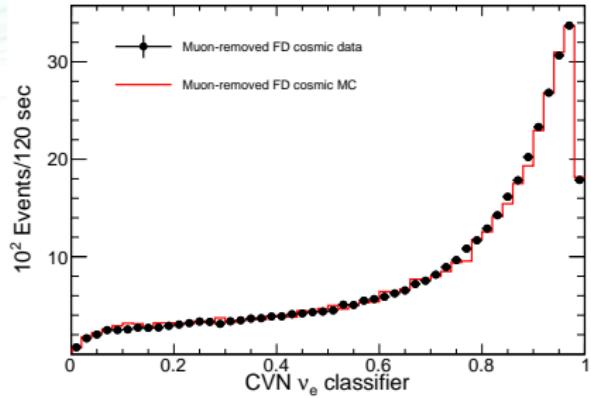
- ▶ Expect 473 FD ν_μ CC events with no oscillation
- ▶ Observe 82 (inc. 3.7 beam bkg. and 2.9 cosmic)

$$\Delta m_{32}^2 = -2.71 \pm 0.12 \times 10^{-3} \text{ eV}^2 \text{ (NH)}$$

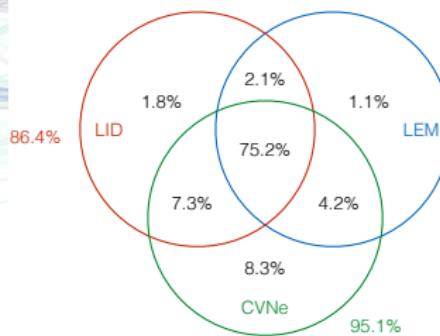
$$\sin^2 \theta_{23} = 0.40^{+0.03}_{-0.02} (0.62^{+0.02}_{-0.03})$$

- ▶ Maximal mixing excluded at 2.5σ (FC corrections in progress)

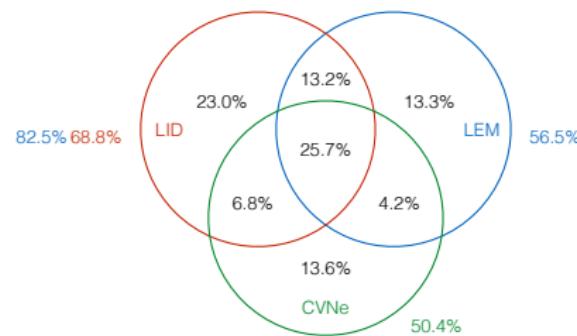
NOvA Preliminary



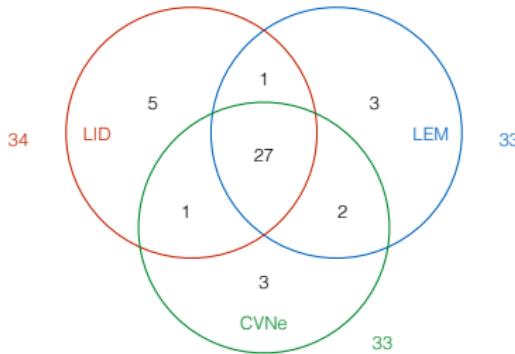
Signal MC



Background MC

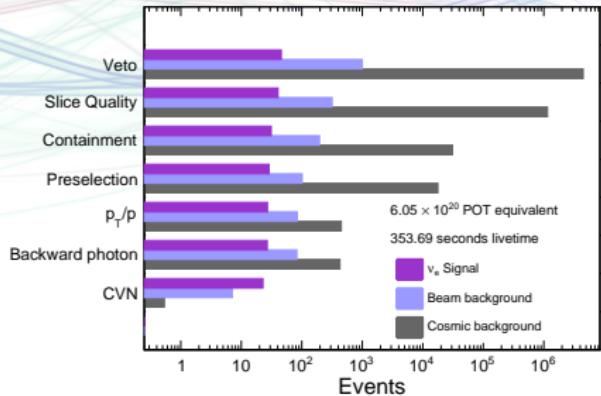


Data

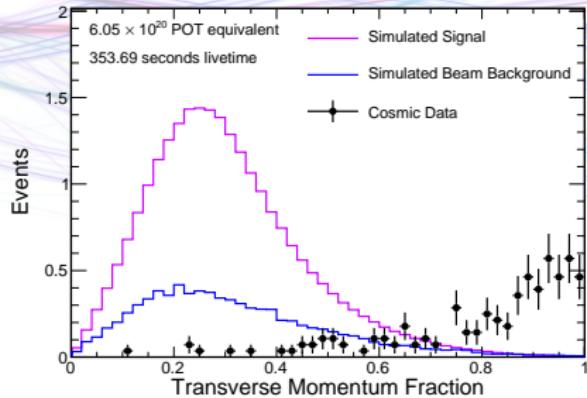


ν_e extrapolation and FD cosmic rejection

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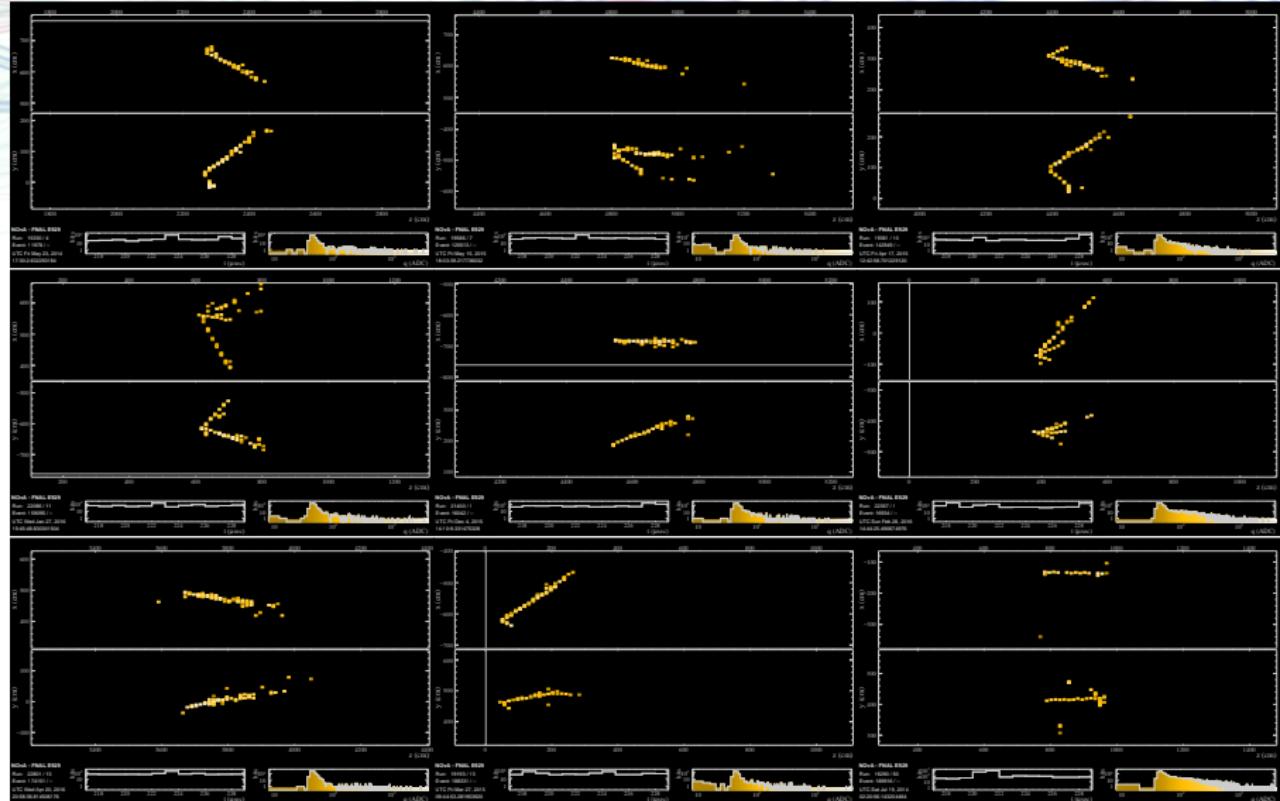


NOvA Preliminary

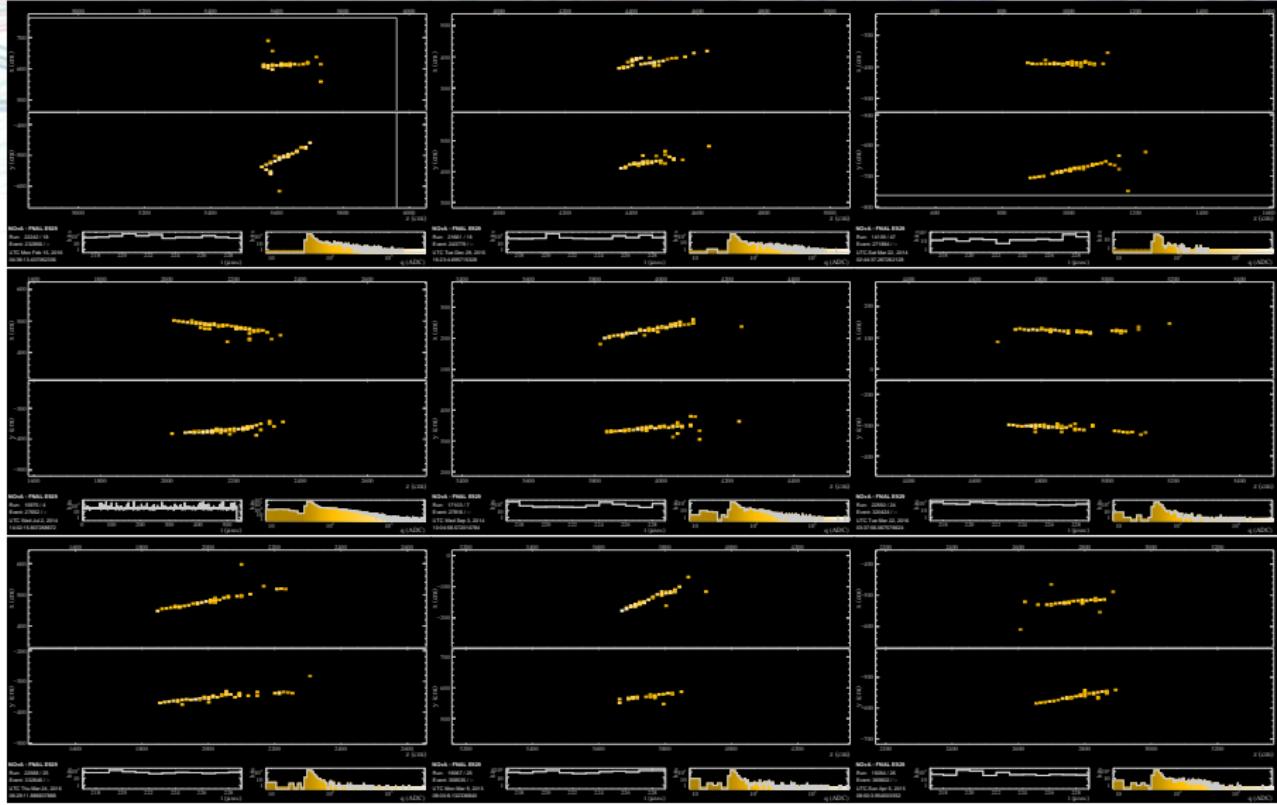


- After decomposition, background components extrapolated by F/N
- Signal comes from oscillated ND ν_μ , extrapolate ν_μ spectrum
- Require containment of candidate electron shower
- Transverse momentum fraction: $p_T/p < 0.65$
- CVN also removes a lot of background
- Final cosmic-ray bkg. from timing sideband: 0.5 events

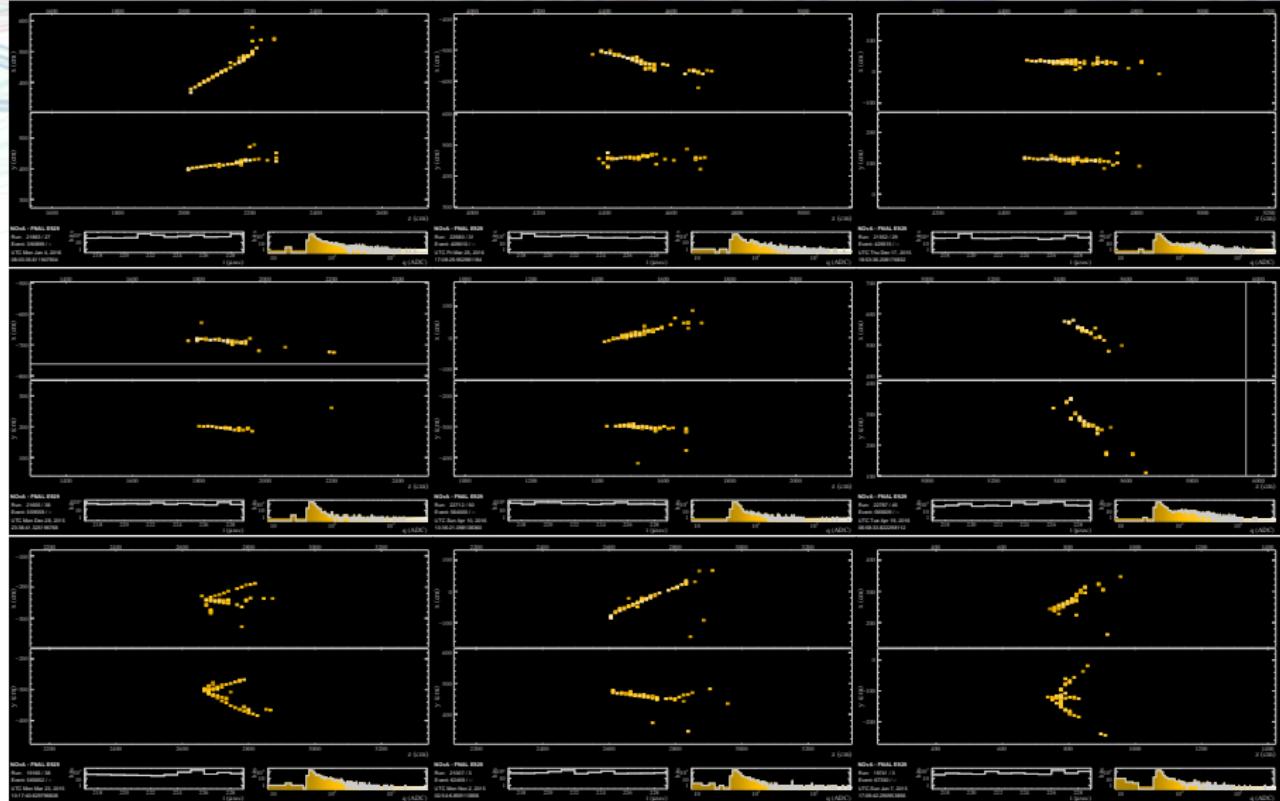
ν_e evds



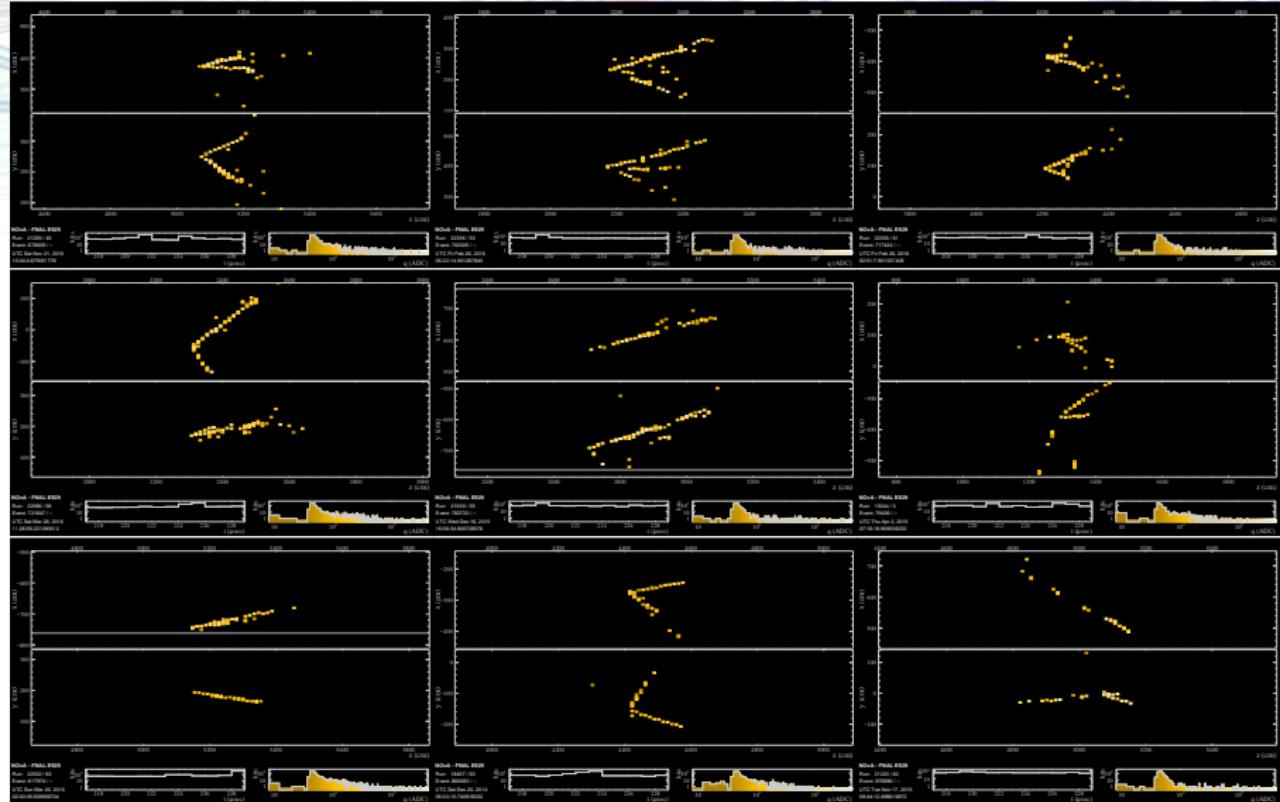
ν_e evds



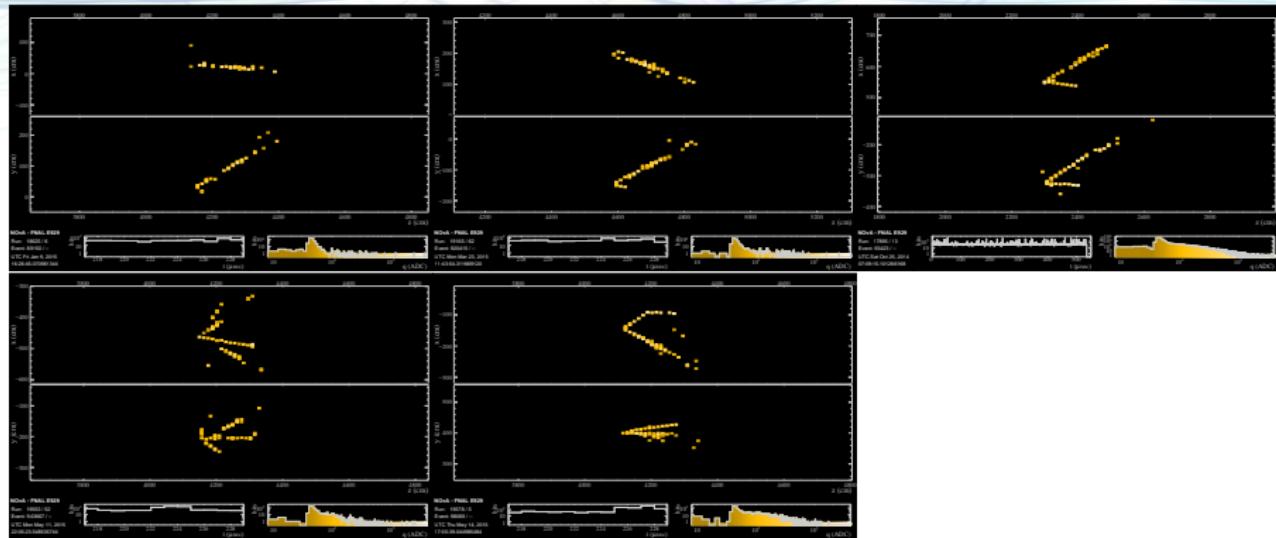
ν_e evds



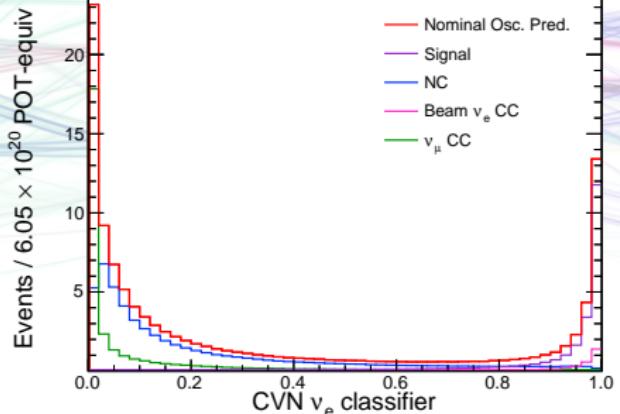
ν_e evds



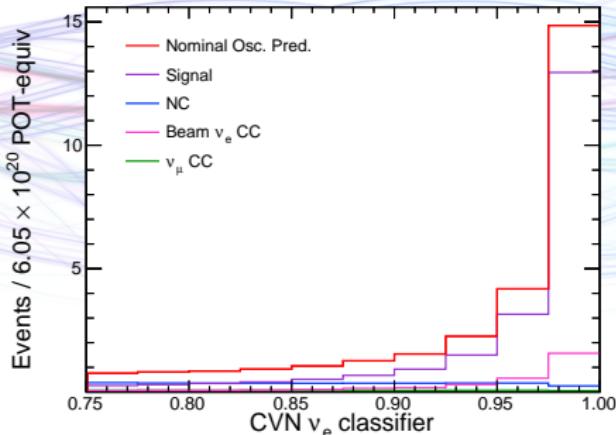
ν_e evds



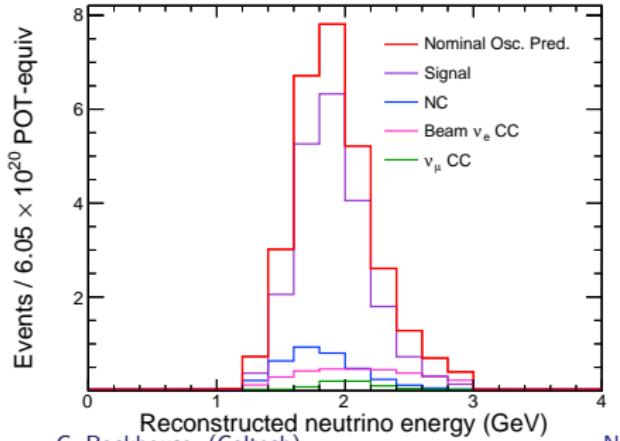
NOvA Simulation



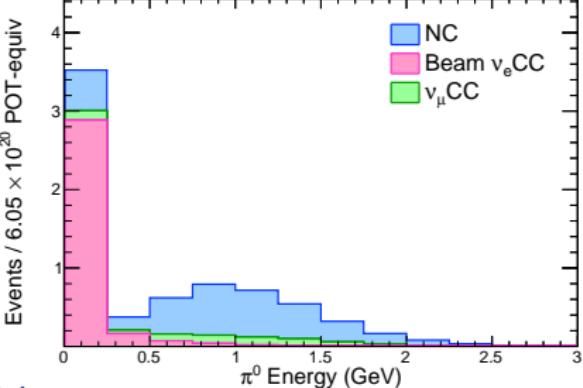
NOvA Simulation



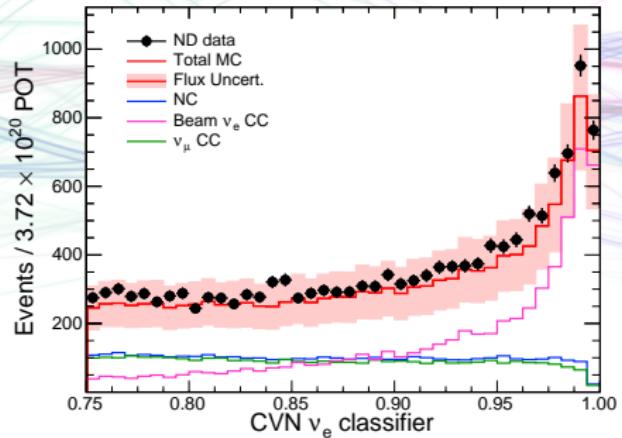
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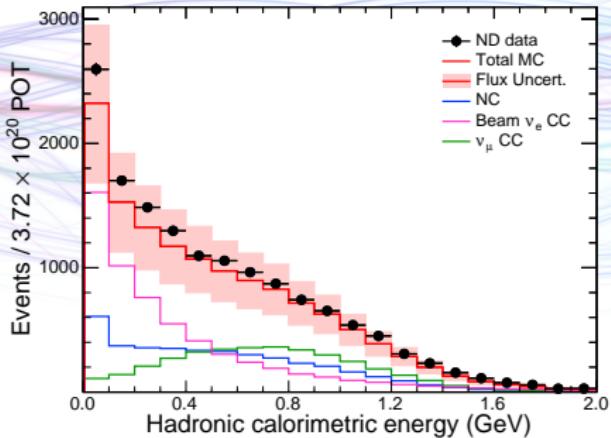
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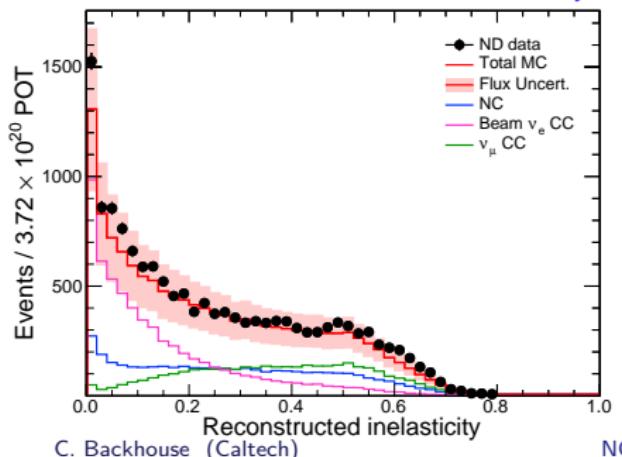
NOvA Preliminary



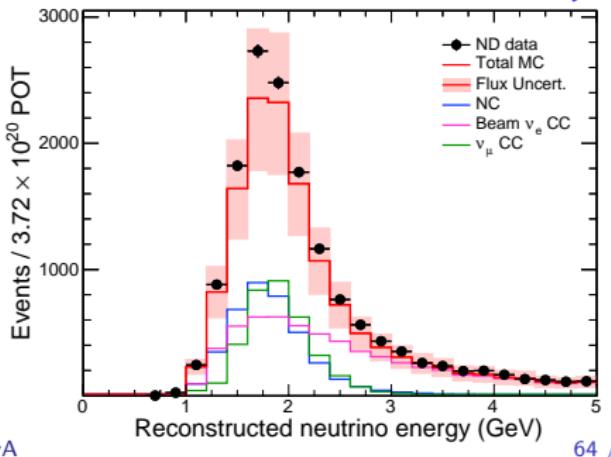
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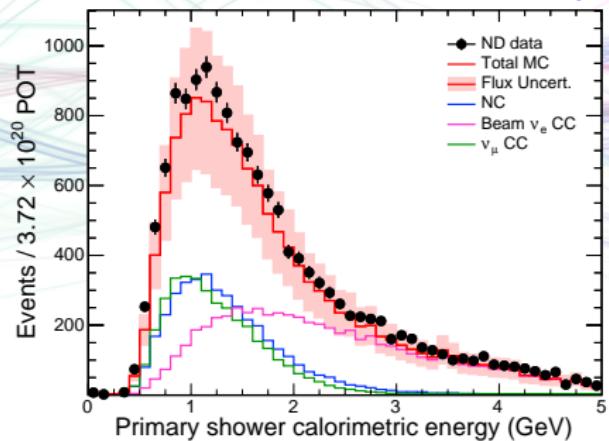
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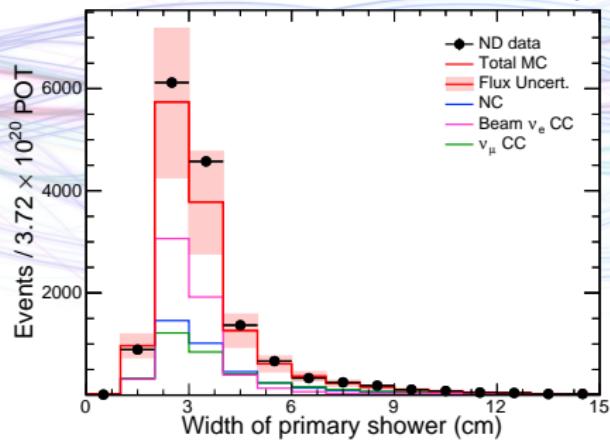
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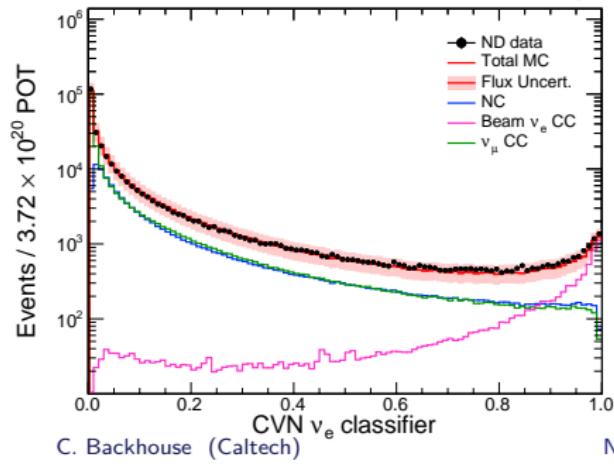
NOvA Preliminary



NOvA Preliminary



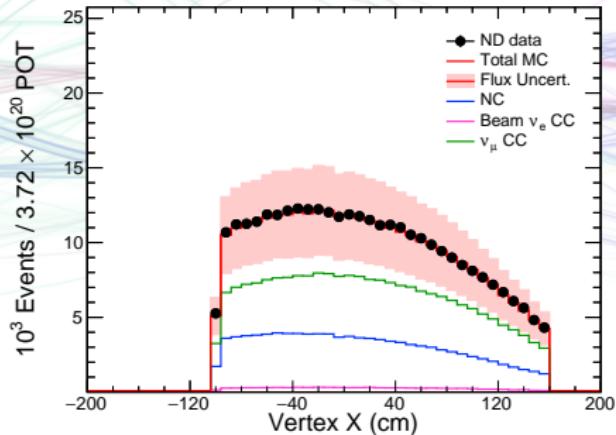
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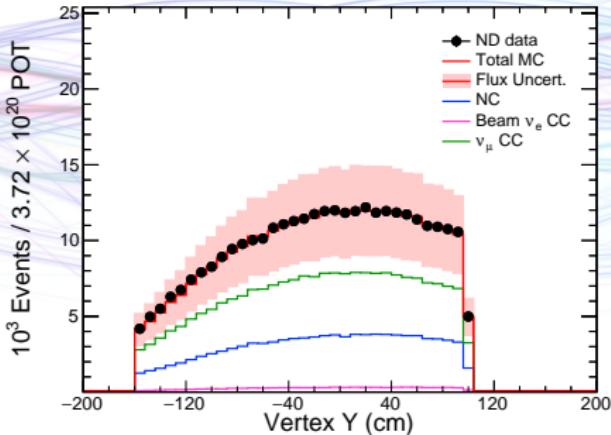
C. Backhouse (Caltech)

NOvA

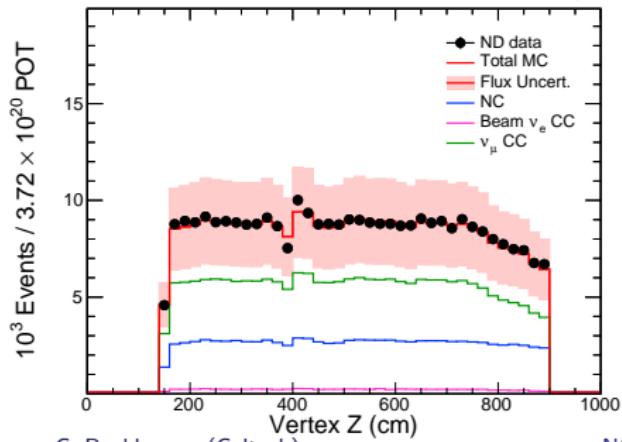
NOvA Preliminary



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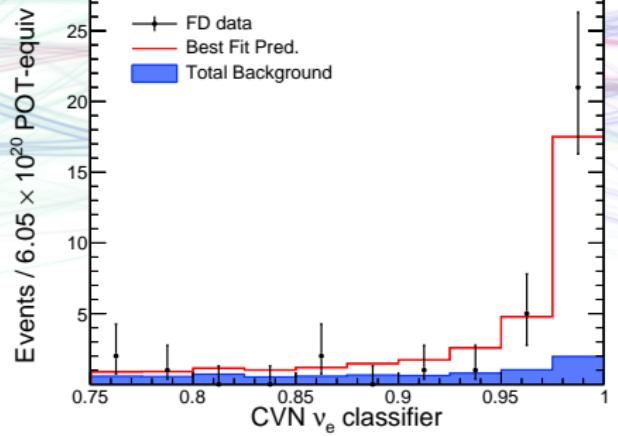
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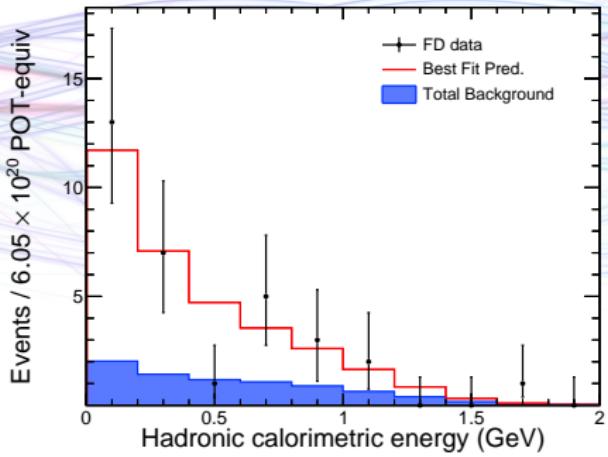
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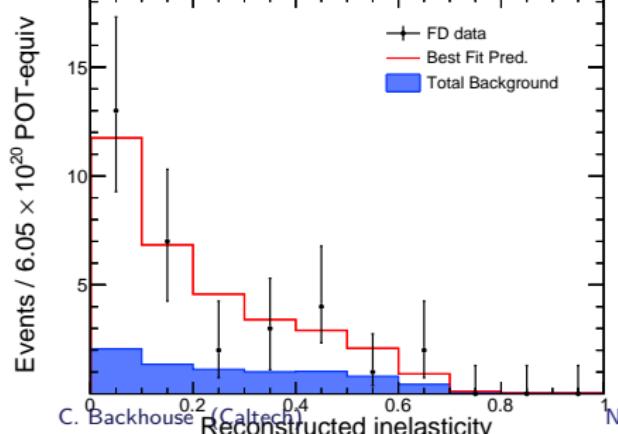
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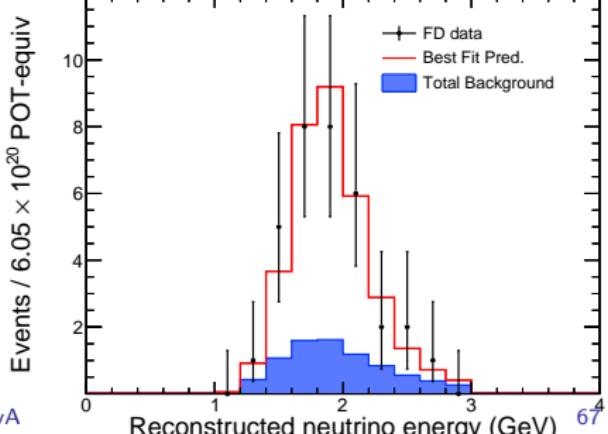
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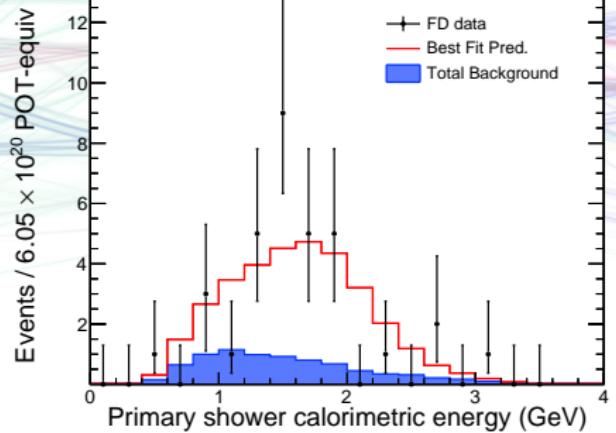
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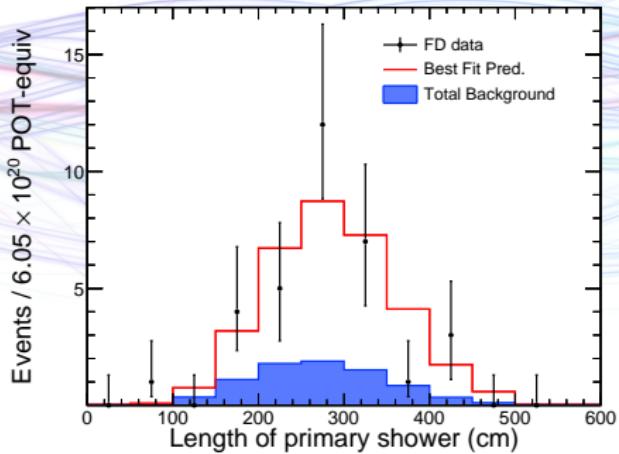
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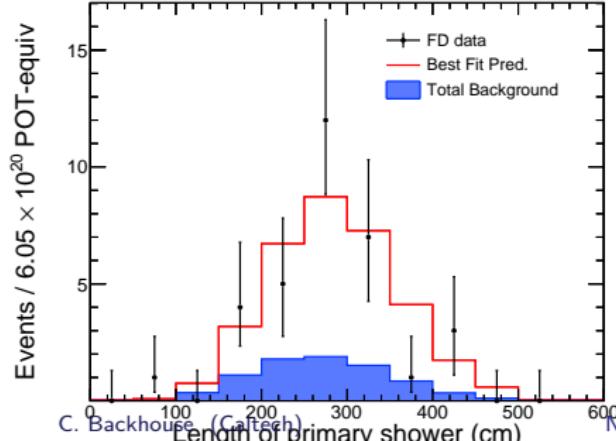
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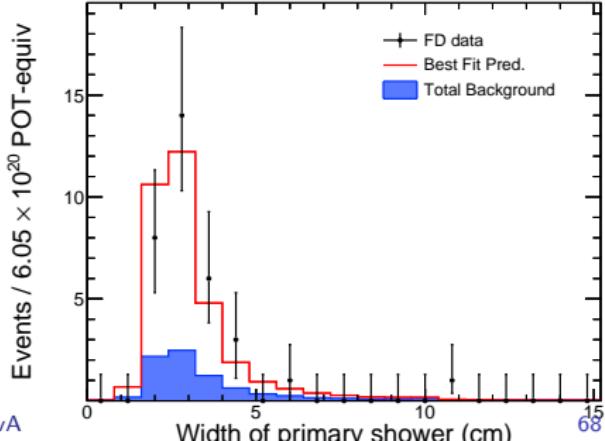
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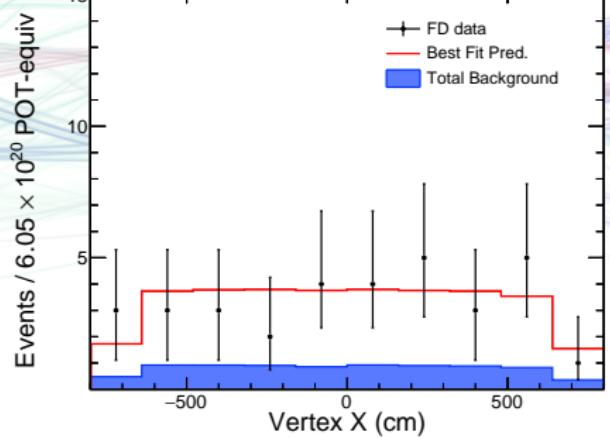
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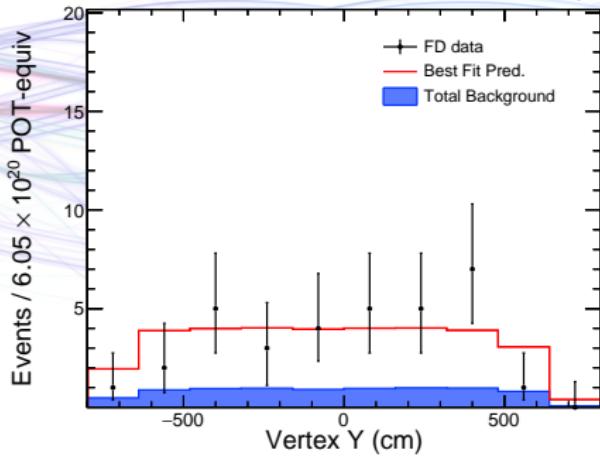
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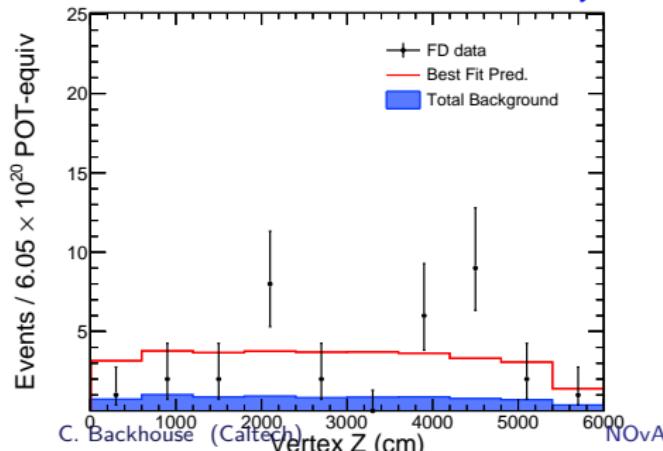
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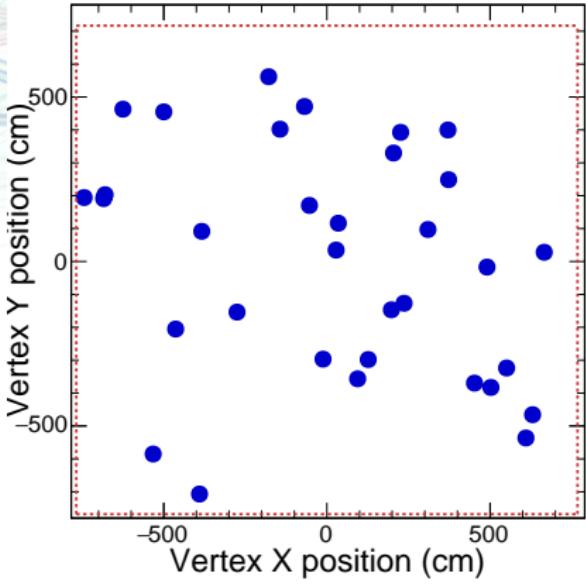
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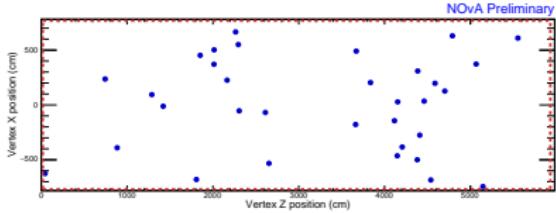
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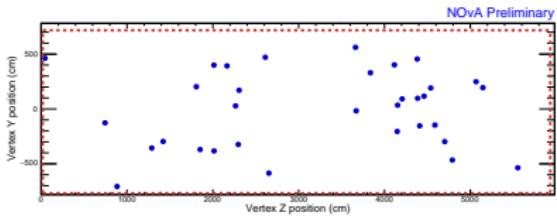
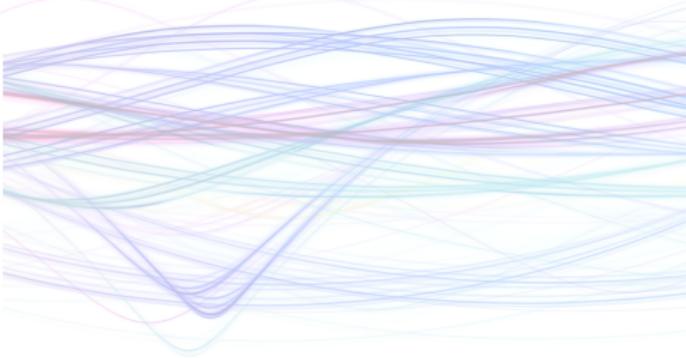
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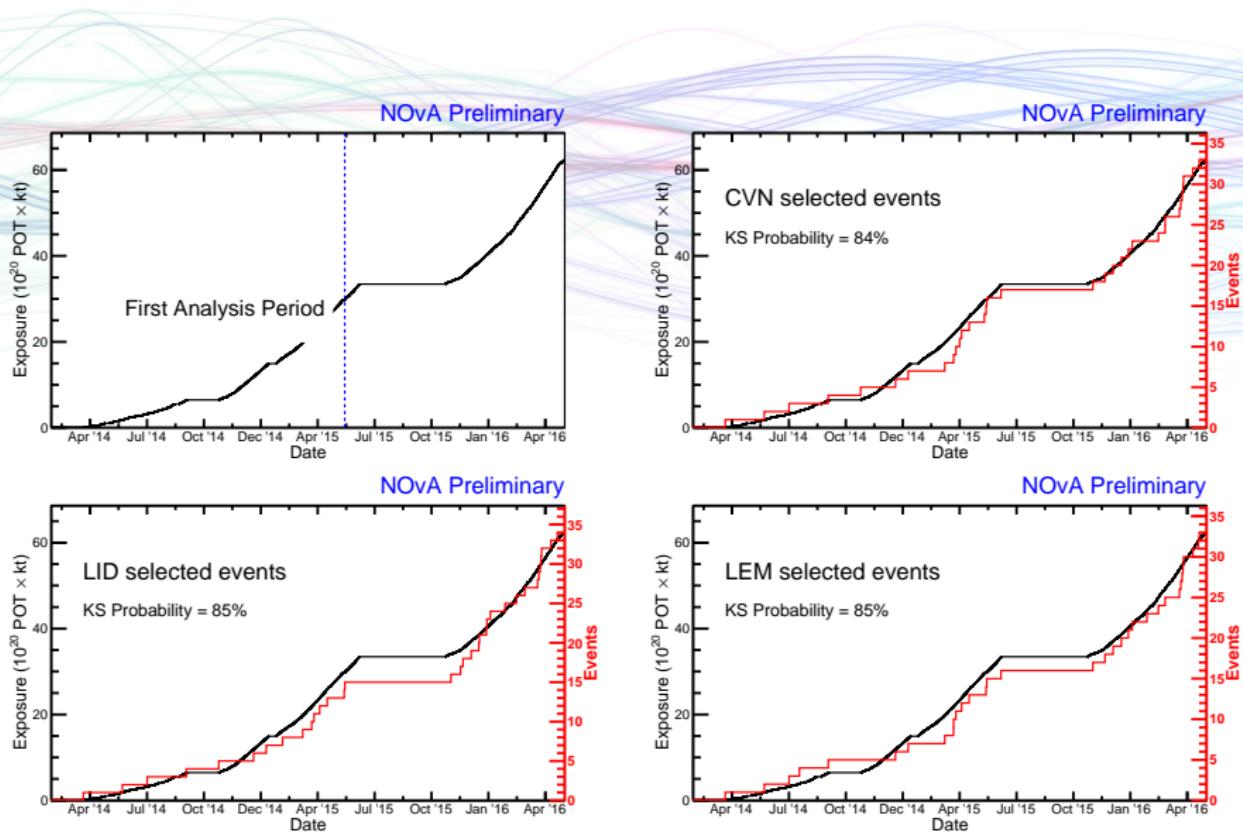
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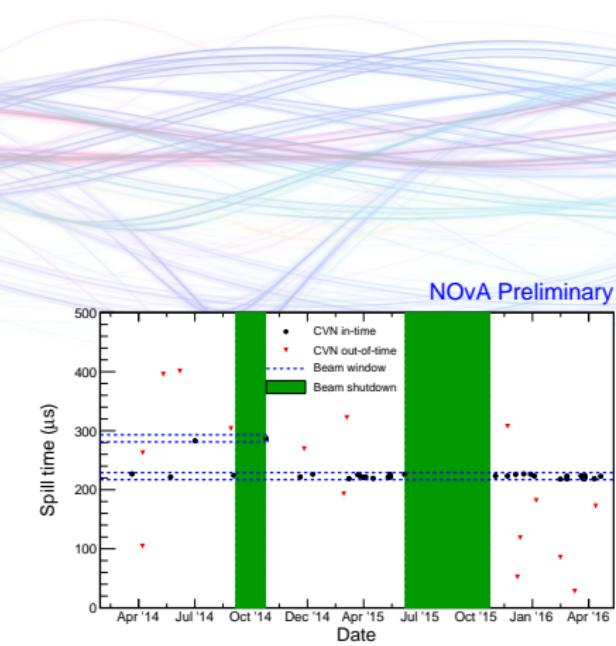
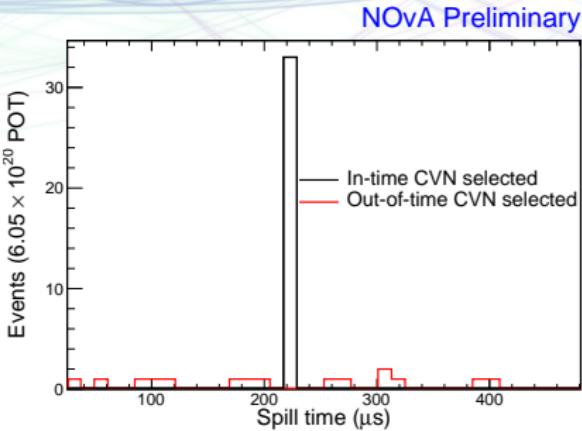


C. Backhouse (Caltech)

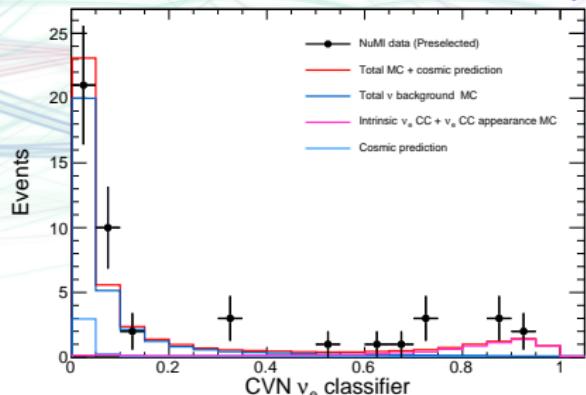


70 / 43

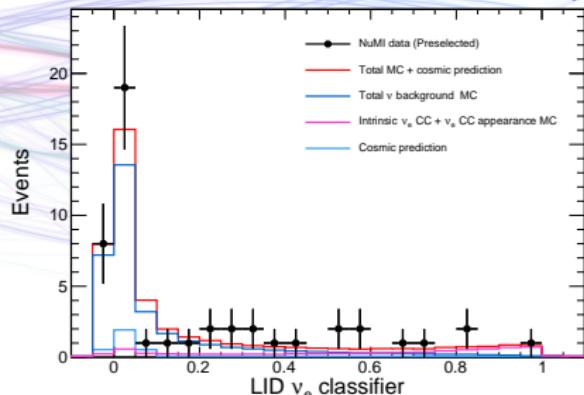




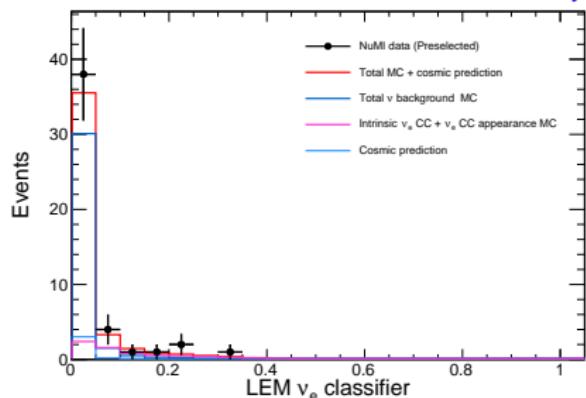
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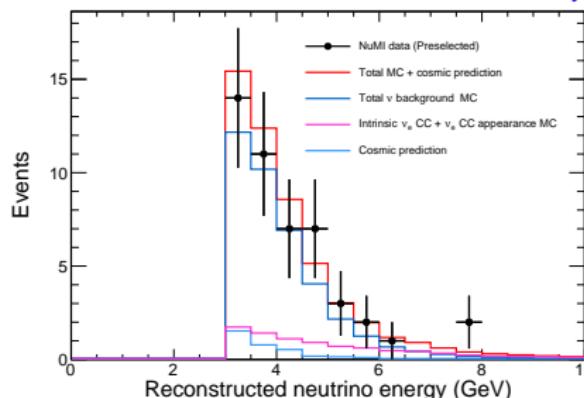
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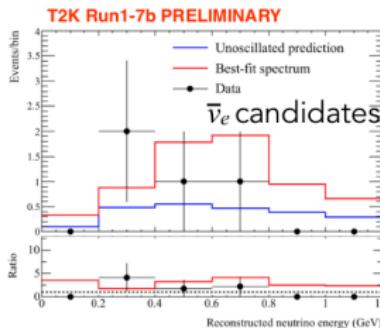
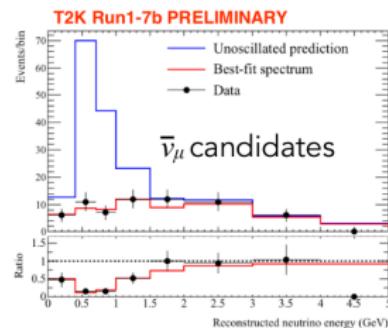
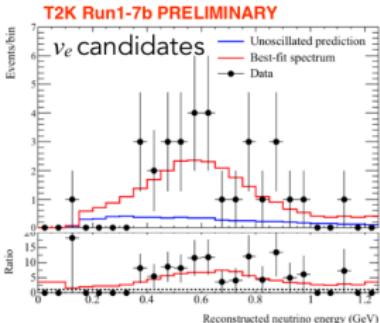
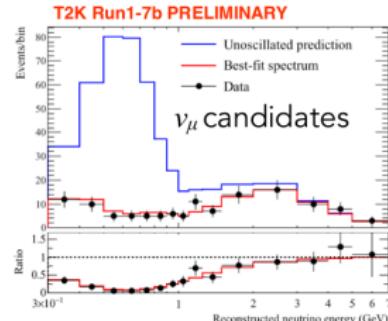
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EVENTS AT SUPER-KAMIOKANDE



E_{rec} distributions assuming 2-body ("QE") kinematics

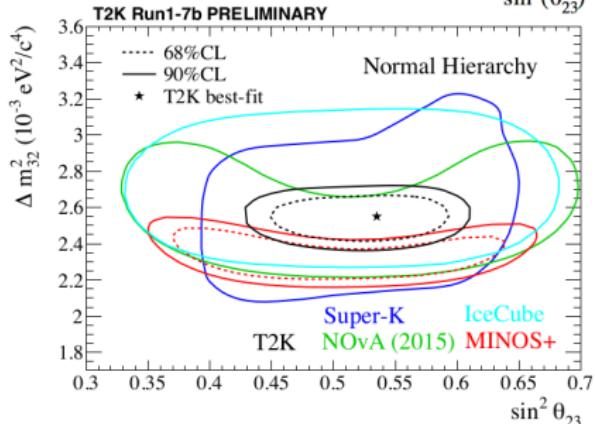
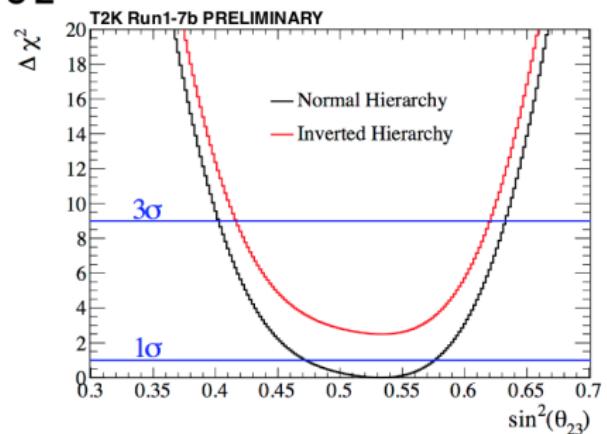
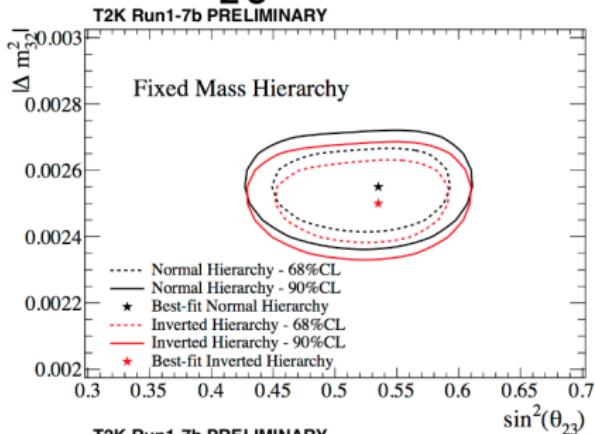
P1.041 R. Shah

	OBS.	EXP. (NH, $\sin^2 \Theta_{23} = 0.528$, NH)			
		$\delta_{CP} = -\pi/2$	$\delta_{CP} = 0$	$\delta_{CP} = +\pi/2$	$\delta_{CP} = \pi$
ν_μ	125	127.9	127.6	127.8	128.1
ν_e	32	27.0	22.7	18.5	22.7
$\bar{\nu}_\mu$	66	64.4	64.3	64.4	64.6
$\bar{\nu}_e$	4	6.0	6.9	7.7	6.8

Analysis frameworks

- Frequentist with $\Delta\chi^2$ fit to
 - $E_{\text{rec}}/\theta_{\text{lep}}$ for $\nu_e/\bar{\nu}_e$
 - E_{rec} for $\nu_\mu/\bar{\nu}_\mu$
- Bayesian with likelihood fit to
 - $p_{\text{lep}}/\theta_{\text{lep}}$ for $\nu_e/\bar{\nu}_e$
 - E_{rec} for $\nu_\mu/\bar{\nu}_\mu$
- Bayesian with Markov Chain MC
 - E_{rec} for all samples
 - simultaneous fit with near detector

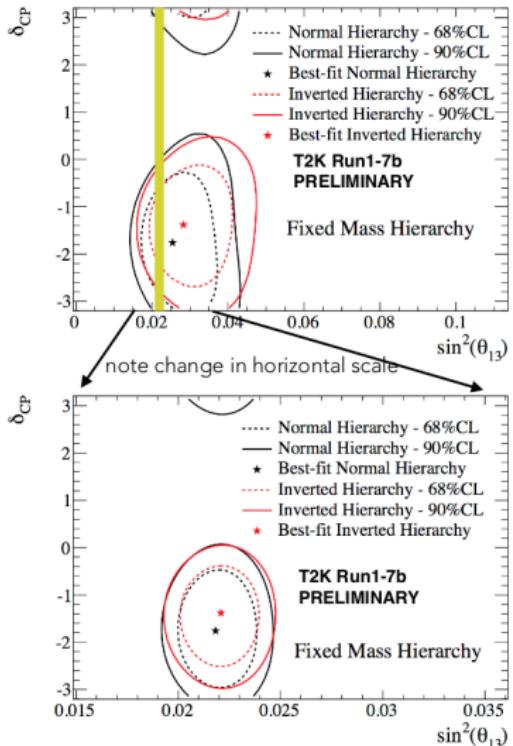
$\sin^2\theta_{23}$ AND Δm^2_{32}



	NH	IH
$\sin^2\theta_{23}$	$0.532^{+0.044}_{-0.060}$	$0.534^{+0.041}_{-0.059}$
$ \Delta m^2_{32} $ ($/10^{-3}\text{eV}^2$)	$2.545^{+0.084}_{-0.082}$	$2.510^{+0.082}_{-0.083}$

- Results continue to be consistent with maximal mixing/oscillation

δ_{CP} VS. θ_{13}



Left: δ_{CP} vs. θ_{13} (fixed $\Delta\chi^2$, fixed hierarchy)

- T2K-only
- T2K with reactor $\sin^2 2\theta_{13} = 0.085 \pm 0.005$

Below: δ_{CP} with Feldman-Cousins critical values and reactor θ_{13}

$$\delta_{CP} = [-3.02, -0.49] \text{ (NH)}, [-1.87, -0.98] \text{ (IH)} @90\% \text{ CL}$$

